

# Neutron production and moderation

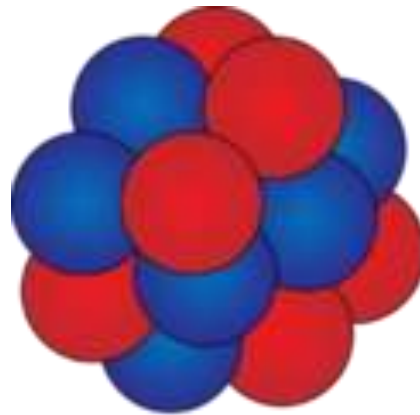
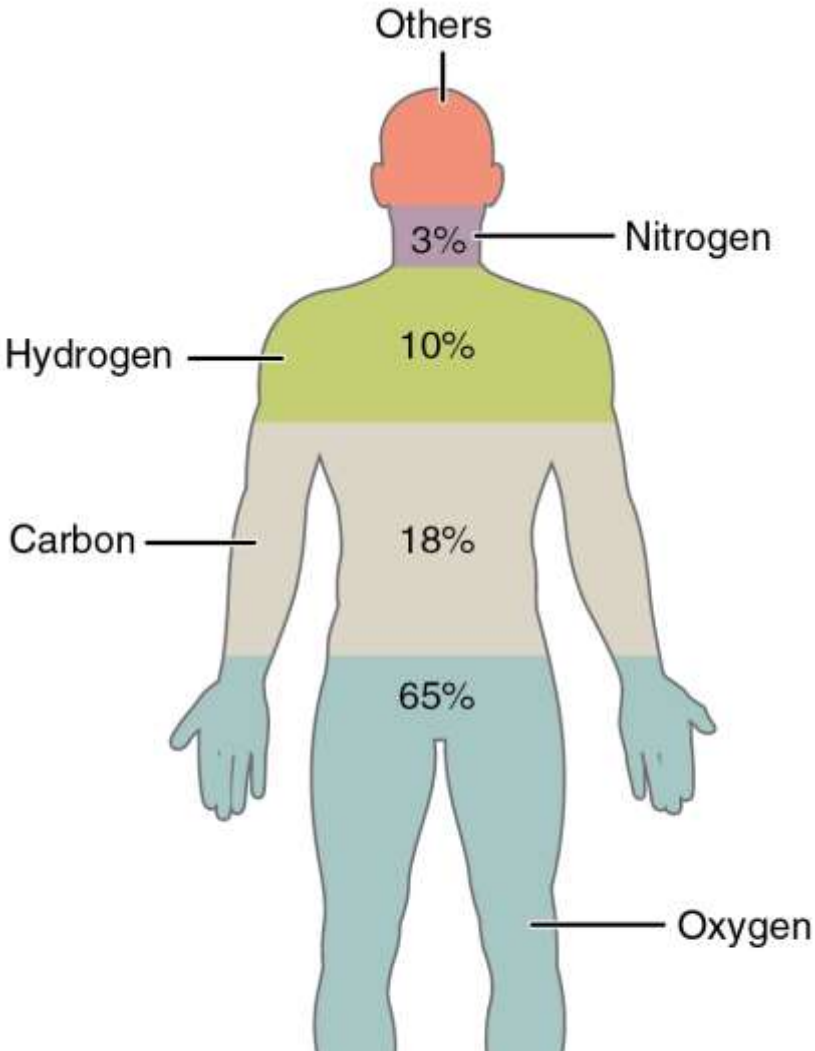
Ulli Köster, ILL

[koester@ill.fr](mailto:koester@ill.fr)

neutrons are everywhere



# Bound neutrons are everywhere

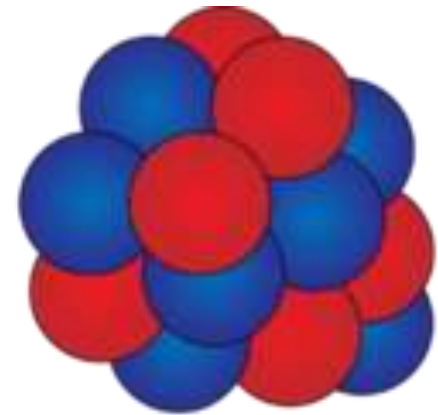


**Carbon-12**

98.9%

6 protons

6 neutrons



**Carbon-13**

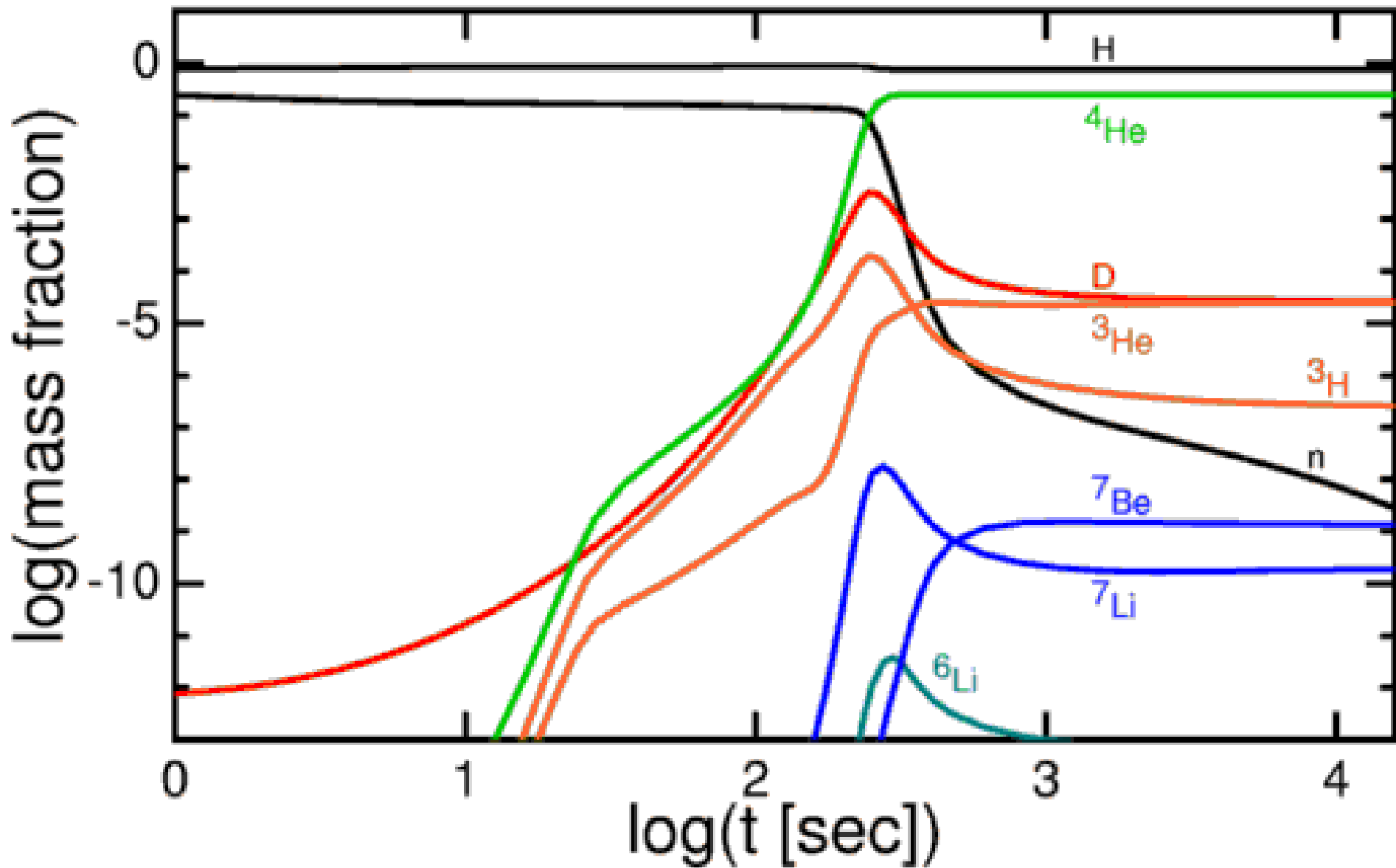
1.1%

6 protons

7 neutrons

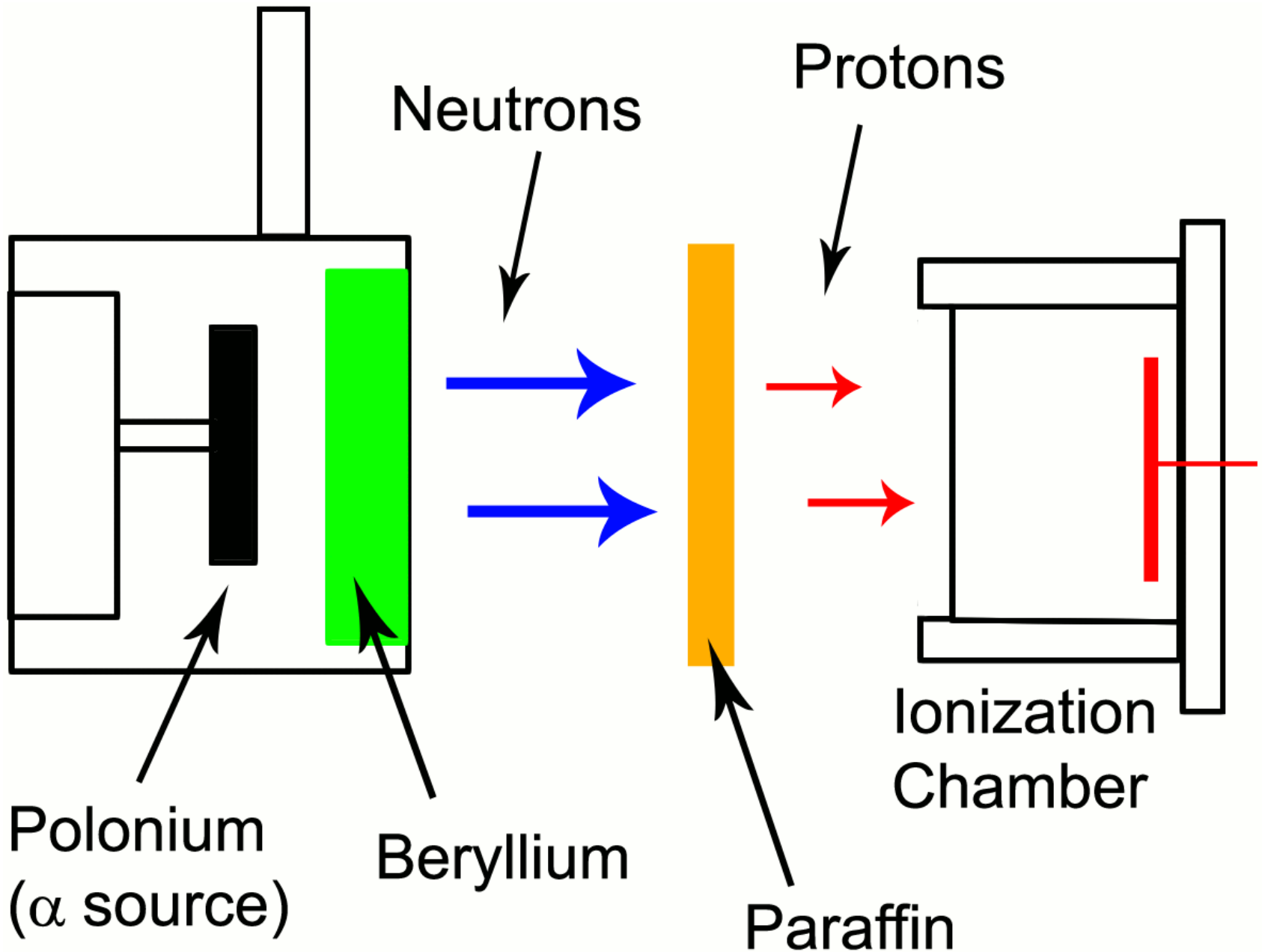
45% neutrons

# Big Bang Nucleosynthesis



Free neutrons have become rare

# Discovery of the neutron



# How neutrons are born

## 1. Alpha-induced reactions: ${}^9\text{Be}(\alpha, n){}^{12}\text{C} + 5.7 \text{ MeV}$

<b>Po 208</b> 2.898 a $\alpha$ 5.1152... $\epsilon$ $\gamma$ (292, 571...) g	<b>Po 209</b> 102 a $\alpha$ 4.881... $\epsilon$ $\gamma$ (895, 261 263...)	<b>Po 210</b> 138.38 d $\alpha$ 5.30438... $\gamma$ (803) $\sigma < 0.0005 + < 0.030$ $\sigma_{n,\alpha} 0.07$ $\sigma_f < 0$
<b>Bi 207</b> 31.55 a $\epsilon, \beta^+ \dots$ $\gamma$ 570, 1064 1770...	<b>Bi 208</b> $3.68 \cdot 10^5 \text{ a}$ $\epsilon$ $\gamma$ 26...	<b>Bi 209</b> 100 $1.9 \cdot 10^{19} \text{ a}$ $\alpha$ 3.077... $\sigma 0.011 + 0.023$ $\sigma_{n,\alpha} < 3E-7$
<b>Pb 206</b> 24.1 $\sigma 0.027$	<b>Pb 207</b> 22.1 $\sigma 0.61$	<b>Pb 208</b> 52.4 $\sigma 0.00023$ $\sigma_{n,\alpha} < 8E-6$

<b>C 11</b> 20.38 m $\beta^+ 1.0$ no $\gamma$	<b>C 12</b> 98.93 $\sigma 0.0035$	<b>C 13</b> 1.07 $\sigma 0.0014$
<b>B 10</b> 19.9 $\sigma 0.3$ $\sigma_{n,\alpha} 3840$ $\sigma_{n,p} 0.007$	<b>B 11</b> 80.1 $\sigma 0.005$	<b>B 12</b> 20.20 ms $\beta^- 13.4 \dots$ $\gamma 4439 \dots$ $\beta\alpha 0.2 \dots$
<b>Be 9</b> 100 $\sigma 0.0078$	<b>Be 10</b> $1.387 \cdot 10^5 \text{ a}$ $\beta^- 0.6$ no $\gamma$ $\sigma < 0.001$	<b>Be 11</b> 13.8 s $\beta^- 11.5 \dots$ $\gamma 2125, 6791 \dots$ $\beta\alpha 0.77, 0.29$

$\alpha$

$\alpha$

n

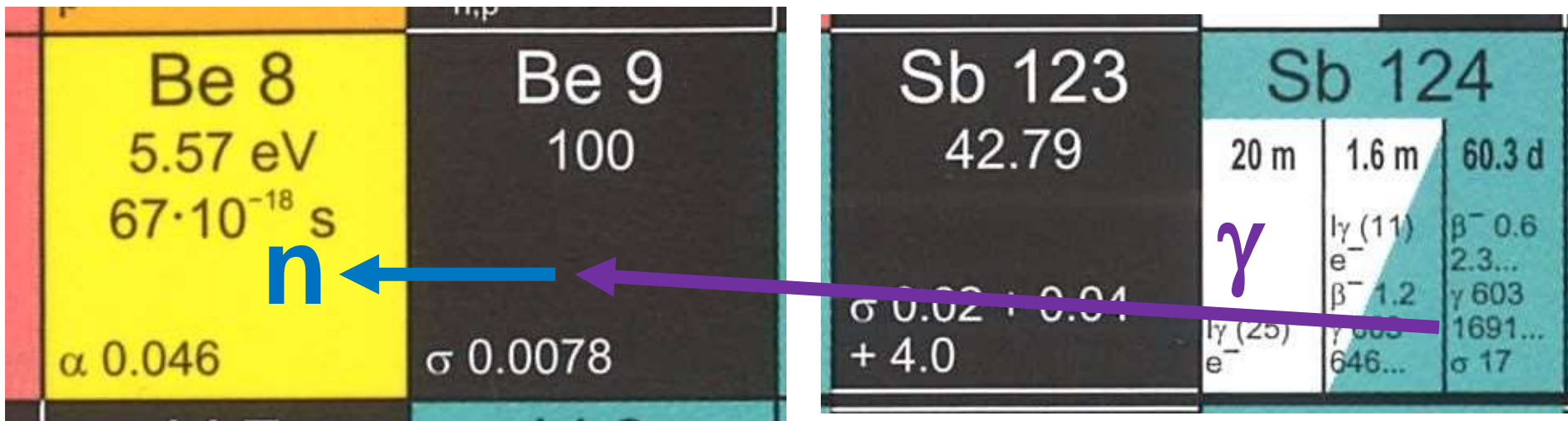
# How neutrons are born

1. Alpha-induced reactions:  ${}^9\text{Be}(\alpha, n){}^{12}\text{C} + 5.7 \text{ MeV}$
2. Deuteron fusion:  $d(d, n){}^3\text{He} + 3.3 \text{ MeV}$ ,  $t(d, n){}^4\text{He} + 17.6 \text{ MeV}$

<b>He</b> 4.002602 $\sigma_{\text{abs}} < 0.05$	<b>He 3</b> 0.000134 $\sigma$ 0.00005 $\sigma_{n,p}$ 5330	<b>He 4</b> 99.999866 $\sigma$ 0.00005 $\sigma_{n,p}$ 5330
<b>H 1</b> 99.9885 $\sigma$ 0.332	<b>H 2</b> 0.0115 $\sigma$ 0.00051	<b>H 3</b> 12.312 a $\beta^-$ 0.0185743 $\sigma < 6E-6$

# How neutrons are born

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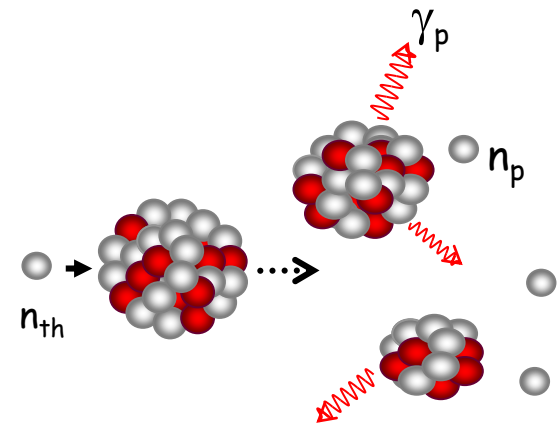
<b>Cf 250</b> 13.08 a $\alpha$ 6.030, 5.989... sf $\gamma$ (43...), $e^-$ $\sigma$ 2000, $\sigma_f$ 110	<b>Cf 251</b> 898 a $\alpha$ 5.679, 5.849 6.012... $\gamma$ 177, 227... $\sigma$ 2900, $\sigma_f$ 4500	<b>Cf 252</b> 2.645 a $\alpha$ 6.118, 6.076... sf $\gamma$ (43...), $e^-$ $\sigma$ 20, $\sigma_f$ 32
<b>Bk 249</b> 330 d $\beta^-$ 0.1 $\alpha$ 5.419, 5.391... sf, $\gamma$ (327, 308...) $\sigma$ 700, $\sigma_f$ -0.1	<b>Bk 250</b> 3.217 h $\beta^-$ 0.7, 1.8... $\gamma$ 989, 1032 1029... $\sigma_f$ 1000	<b>Bk 251</b> 55.6 m $\beta^-$ 0.9, 1.1... $\gamma$ 118, 130 133...
<b>Cm 248</b> $3.40 \cdot 10^5$ a $\alpha$ 5.078, 5.035... sf, $\gamma$ , $e^-$ , g $\sigma$ 2.6, $\sigma_f$ 0.36	<b>Cm 249</b> 64.15 m $\beta^-$ 0.9 $\gamma$ 631, (560 36...), $e^-$ -1.6	<b>Cm 250</b> -9700 a sf $\alpha?$ , $\beta^-?$ $\sigma$ -80



# How neutrons are born

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<b>Pu 237</b> 45.2 d sf $\alpha$ 5.334... $\gamma$ 60... $e^-$ $\sigma$ 2300	<b>Pu 238</b> 87.74 a sf $\alpha$ 5.499, 5.456... Si, Mg $\gamma$ (43, 100...), $e^-$ $\sigma$ 510, $\sigma_f$ 17	<b>Pu 239</b> $2.411 \cdot 10^4$ a sf $\alpha$ 5.157, 5.144... $\gamma$ (52...), $e^-$ m $\sigma$ 270, $\sigma_f$ 752	<b>Pu 240</b> 6563 a sf $\alpha$ 5.166, 5.124... $\gamma$ (45...), $e^-$ g $\sigma$ 290, $\sigma_f$ -0.059	<b>Pu 241</b> 14.35 a sf $\beta^-$ 0.02, g $\alpha$ 4.895... $\gamma$ (146...), $e^-$ $\sigma$ 370, $\sigma_f$ 1010
<b>Np 236</b> 22.5 h $\beta^-$ 0.5... $\gamma$ (642...), $e^-$ g $\sigma$ 2700	<b>Np 237</b> $2.144 \cdot 10^5$ a sf $\alpha$ 4.790, 4.774... $\gamma$ 29, 87... $e^-$ $\sigma$ 170, $\sigma_f$ 0.020	<b>Np 238</b> 2.117 d $\beta^-$ 1.2... $\gamma$ 984 1029, 1026, 924... $e^-$ , g $\sigma_f$ 2600	<b>Np 239</b> 2.355 d $\beta^-$ 0.4, 0.7... $\gamma$ 106, 278 228... $e^-$ , g $\sigma$ 32 + 19, $\sigma_f$ < 1	<b>Np 240</b> 7.22 m $\beta^-$ 2.2... $\gamma$ 555... 597... $e^-$ g
<b>U 235</b> 0.7204 a 26 m $\beta^-$ (0.07) $e^-$	<b>U 236</b> 2.342 $\cdot 10^8$ a 120 ns $\beta^-$ 17.83... $\gamma$ 17.83... 11... $e^-$	<b>U 237</b> 6.75 d $\beta^-$ 0.2... $\gamma$ 60, 208... $e^-$ $\sigma$ ~100 $\sigma_f$ < 0.35	<b>U 238</b> 99.2742 a 298 ns $\beta^-$ 4.468 $\cdot 10^9$ a $\alpha$ 4.198... $\gamma$ 2514... 1879... $e^-$ , $\sigma$ 2.7 $\sigma_f$ 3E-6	<b>U 239</b> 23.5 m $\beta^-$ 1.2, 1.3... $\gamma$ 75, 44... $\sigma$ 22 $\sigma_f$ 15



# How neutrons are born

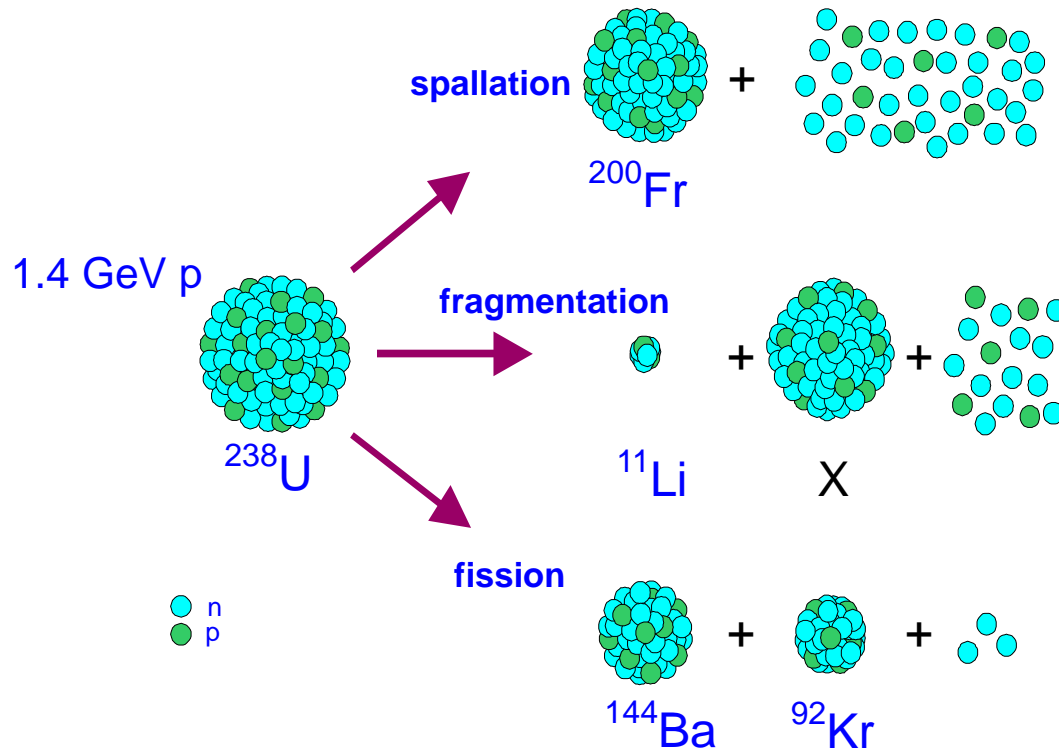
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6. Beta-delayed n emission:  ${}^{87}\text{Br}(\beta^-){}^{87}\text{Kr}^* \rightarrow {}^{86}\text{Kr} + n + 1.3 \text{ MeV}$

Kr 86 17.279 n α 0.003	Kr 87 76.3 m β <sup>-</sup> 3.5, ... γ 403, 255, ... 845...	Kr 88 2.84 h β <sup>-</sup> 0.5, 2.9... γ 2392, 196 2196, 835 1530...
Br 85 2.87 m β <sup>-</sup> 2.5... γ 802, 925... m	Br 86 55.1 s β <sup>-</sup> 3.3, 7.6... γ 1565, 2751...	Br 87 55.7 s β <sup>-</sup> 6.8... γ 1420, 1476, 1578 532, 3000... βn 0.02, 0.85...

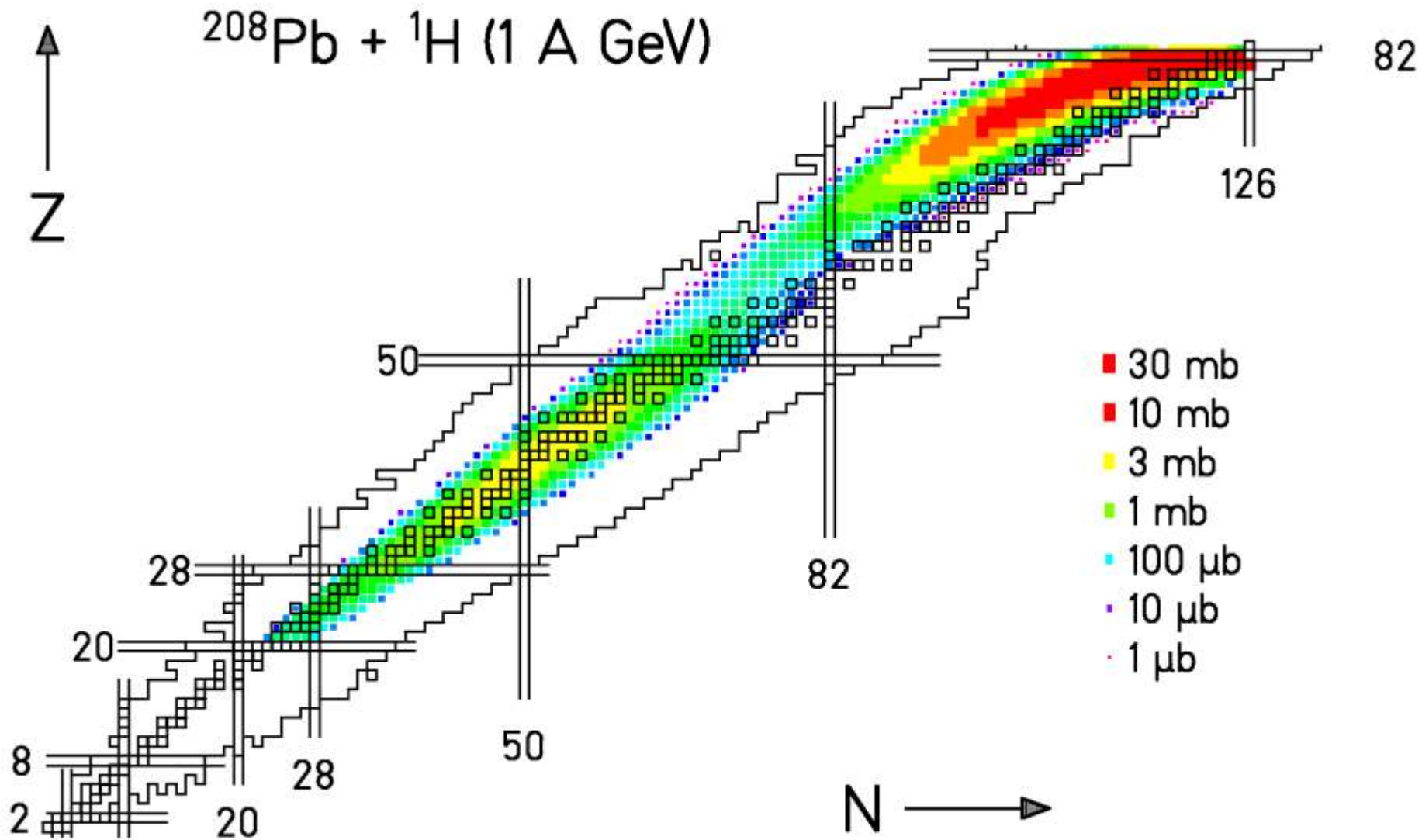
# How neutrons are born

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7. **Spallation:**  ${}^{208}\text{Pb}(p, 3p \ 20n){}^{185}\text{Au} - 173 \text{ MeV}$

# High energy nuclear reactions



# Spallation + Fragmentation + Fission



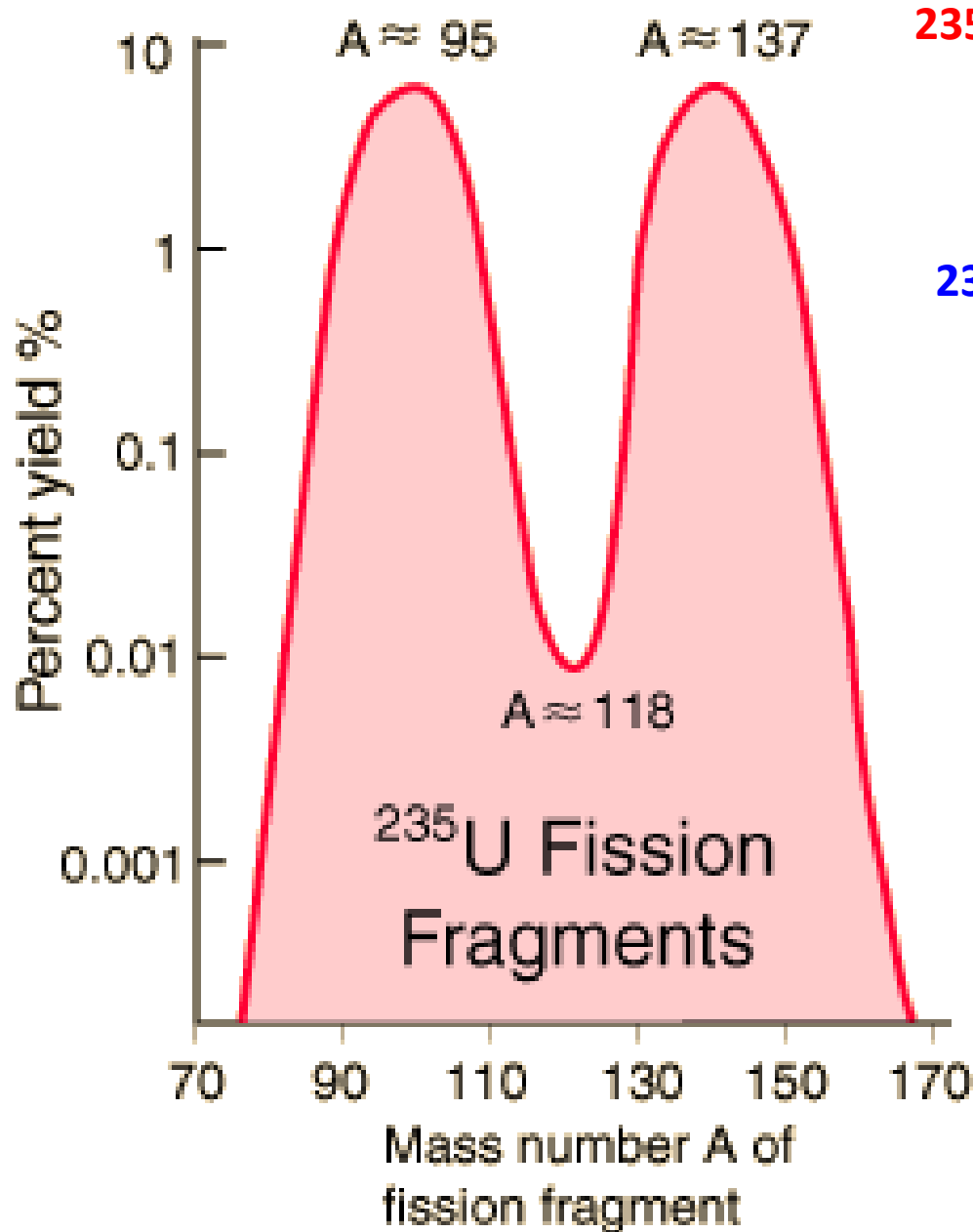
*W. Wlazio et al., Phys. Rev. Lett. 84 (2000) 5736.*

*T. Enqvist et al., Nucl. Phys. A 686 (2001) 481.*

# How neutrons are born

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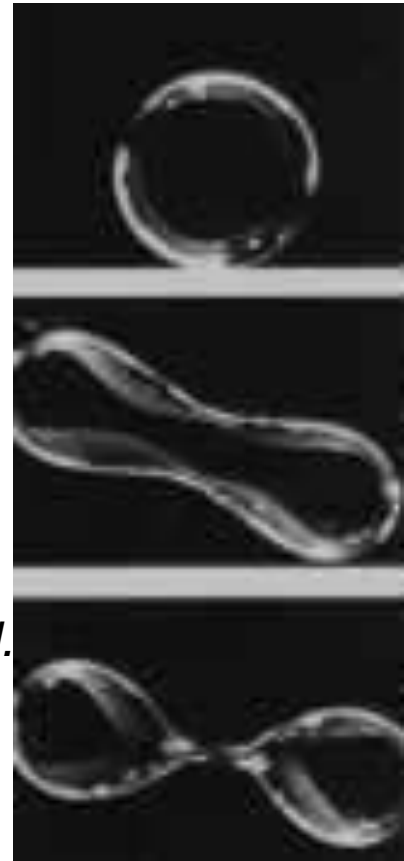
# Mass yields of fission fragments



$$Q_{\text{LDM}} = +184 \text{ MeV}$$



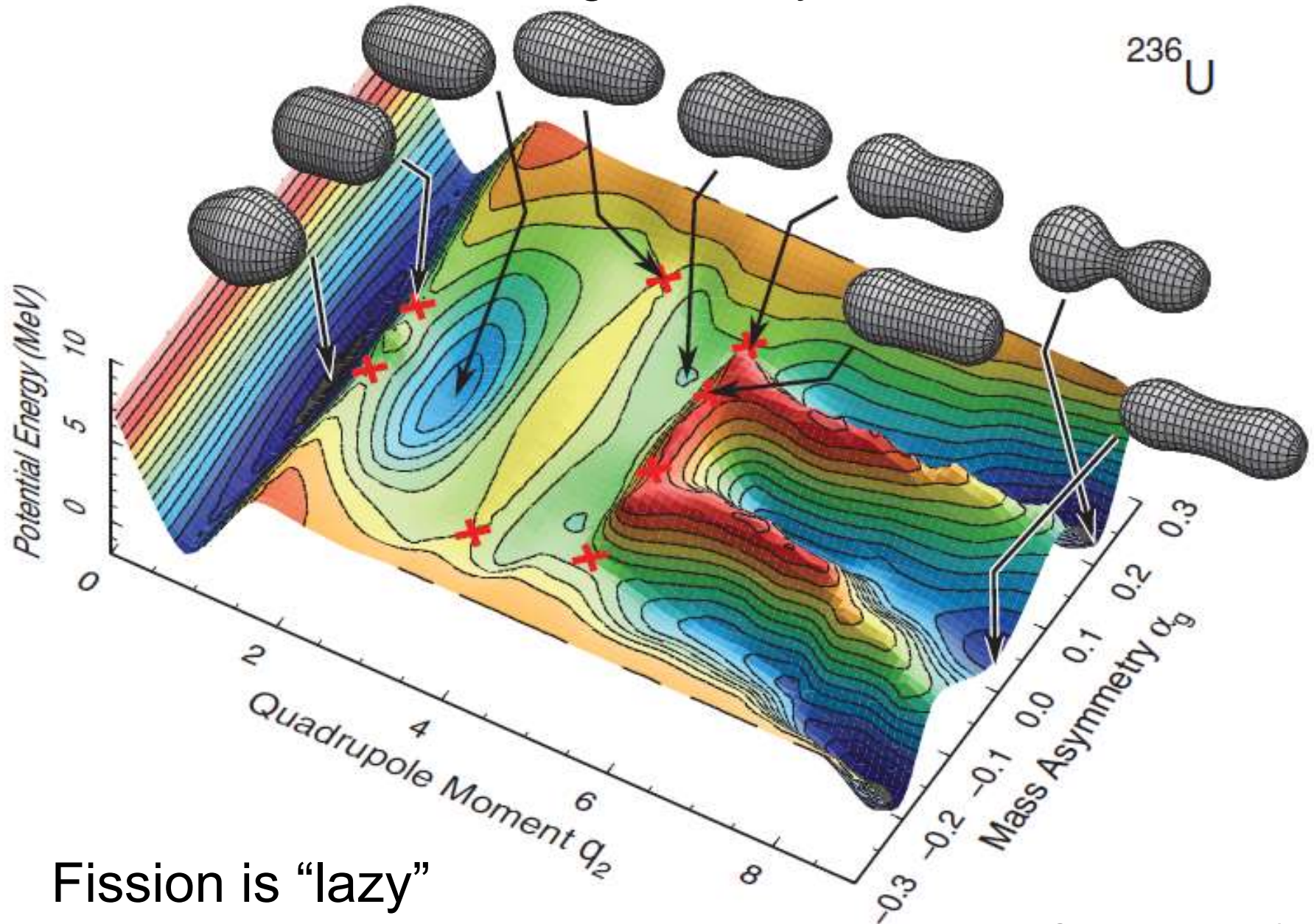
$$Q_{\text{LDM}} = +176 \text{ MeV}$$



*T.G. Wang et al.  
J Fluid Mech  
354, 43 (1994)  
STS-50 flight*

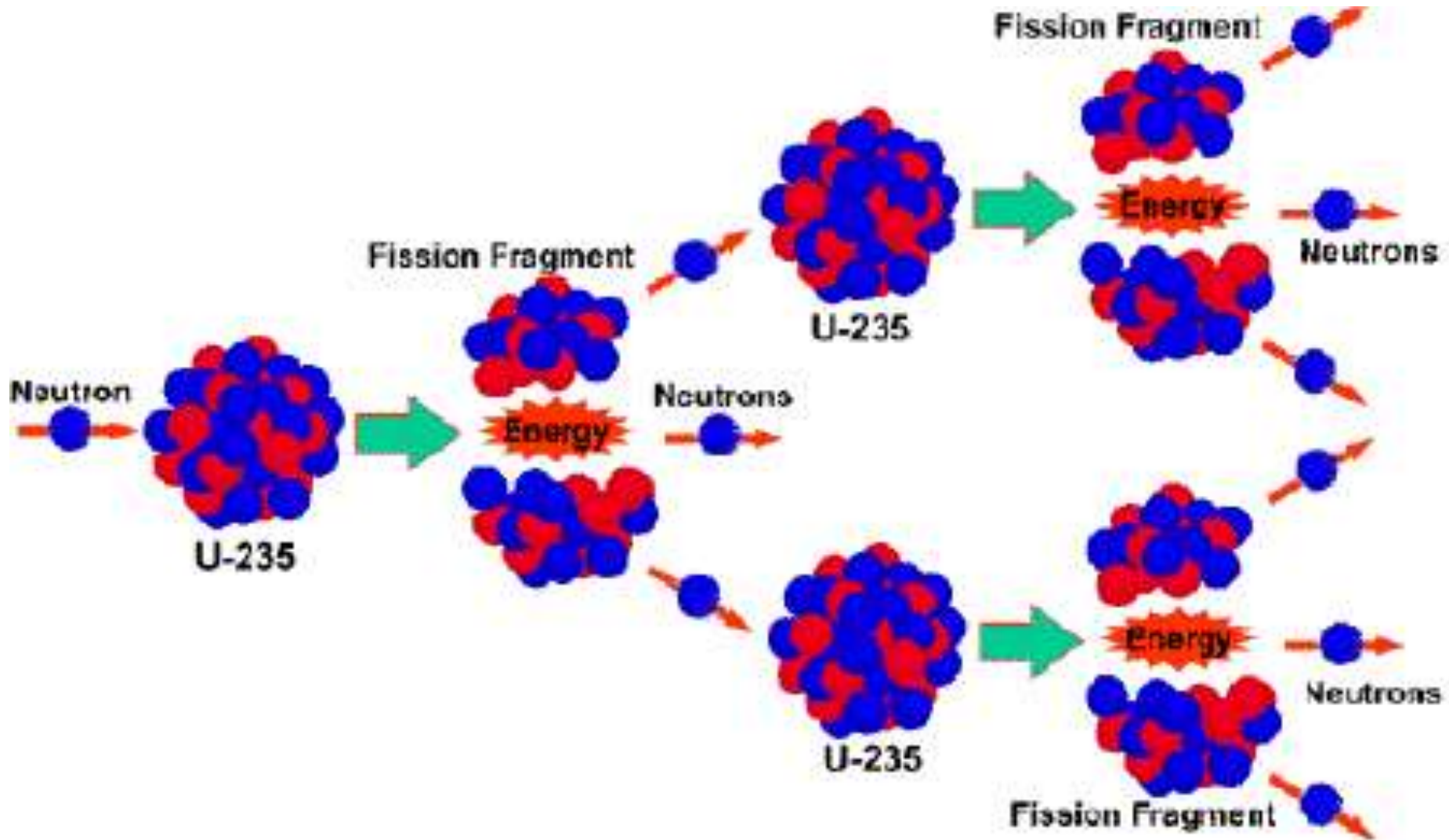


# Understanding fission yields of $^{236}\text{U}$

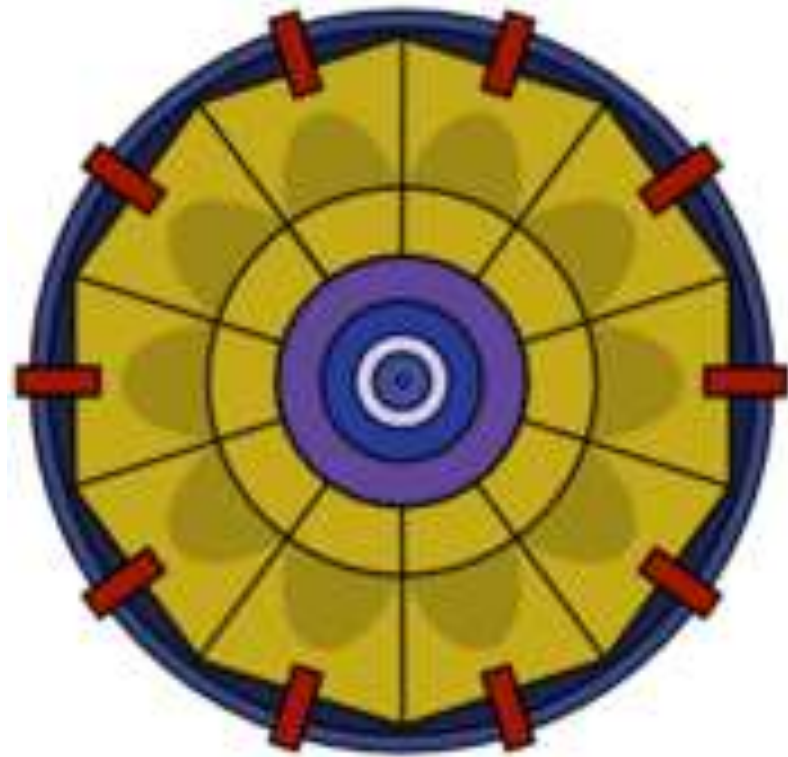


Fission is “lazy”

# A nuclear chain reaction

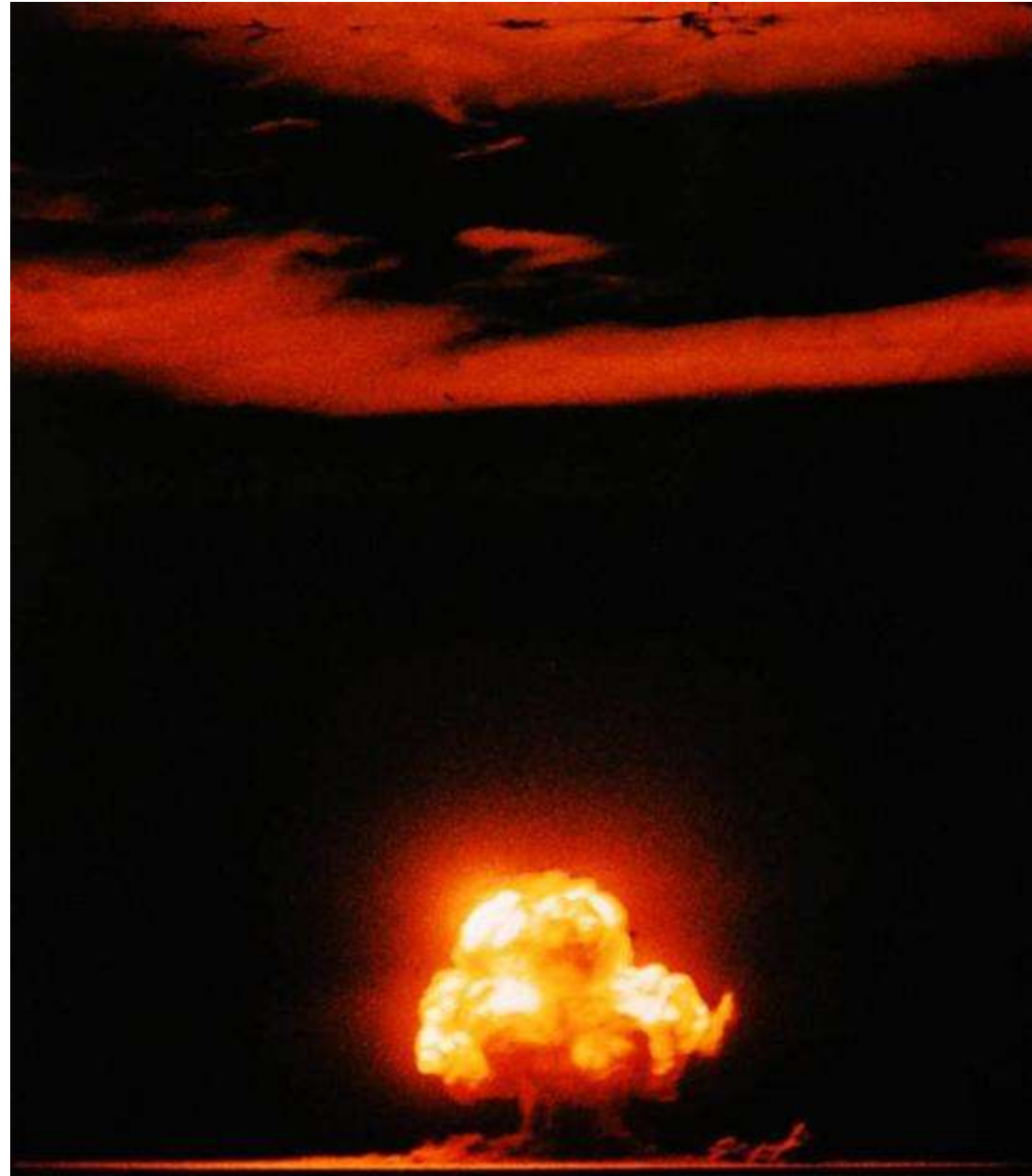


# A single-pulse neutron source

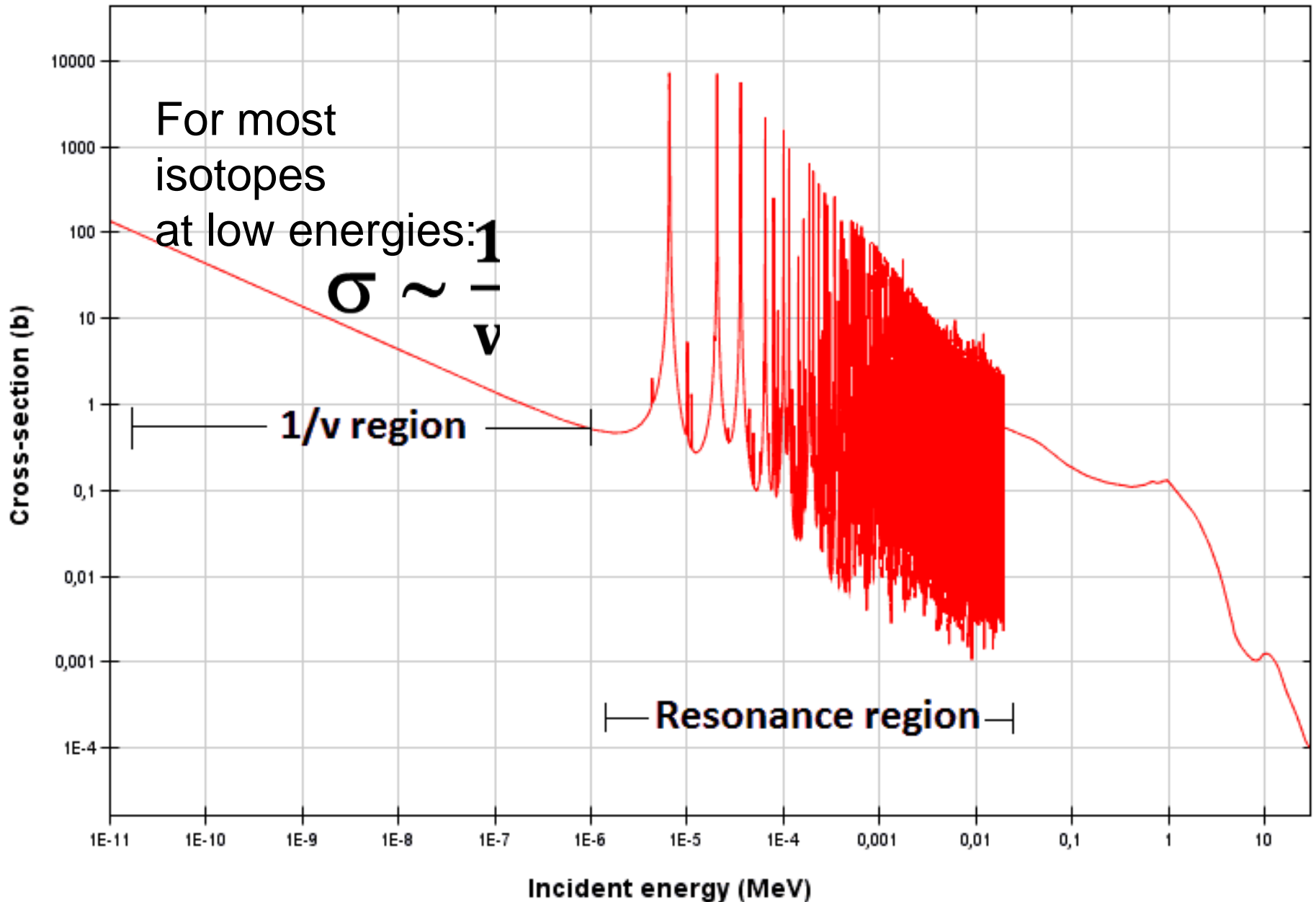


Uncontrolled  
chain reaction  
of fast-neutron  
induced fission

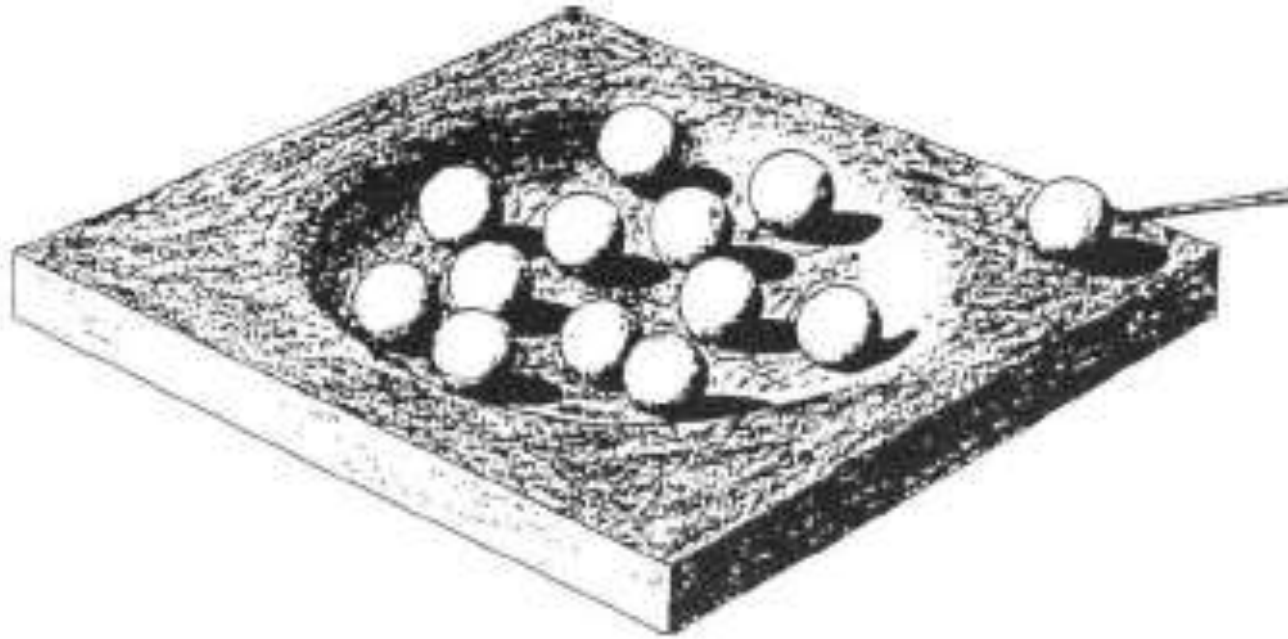
≈25 kg of 93%  $^{235}\text{U}$



# Energy dependence of cross-section



# Bohr's compound nucleus model

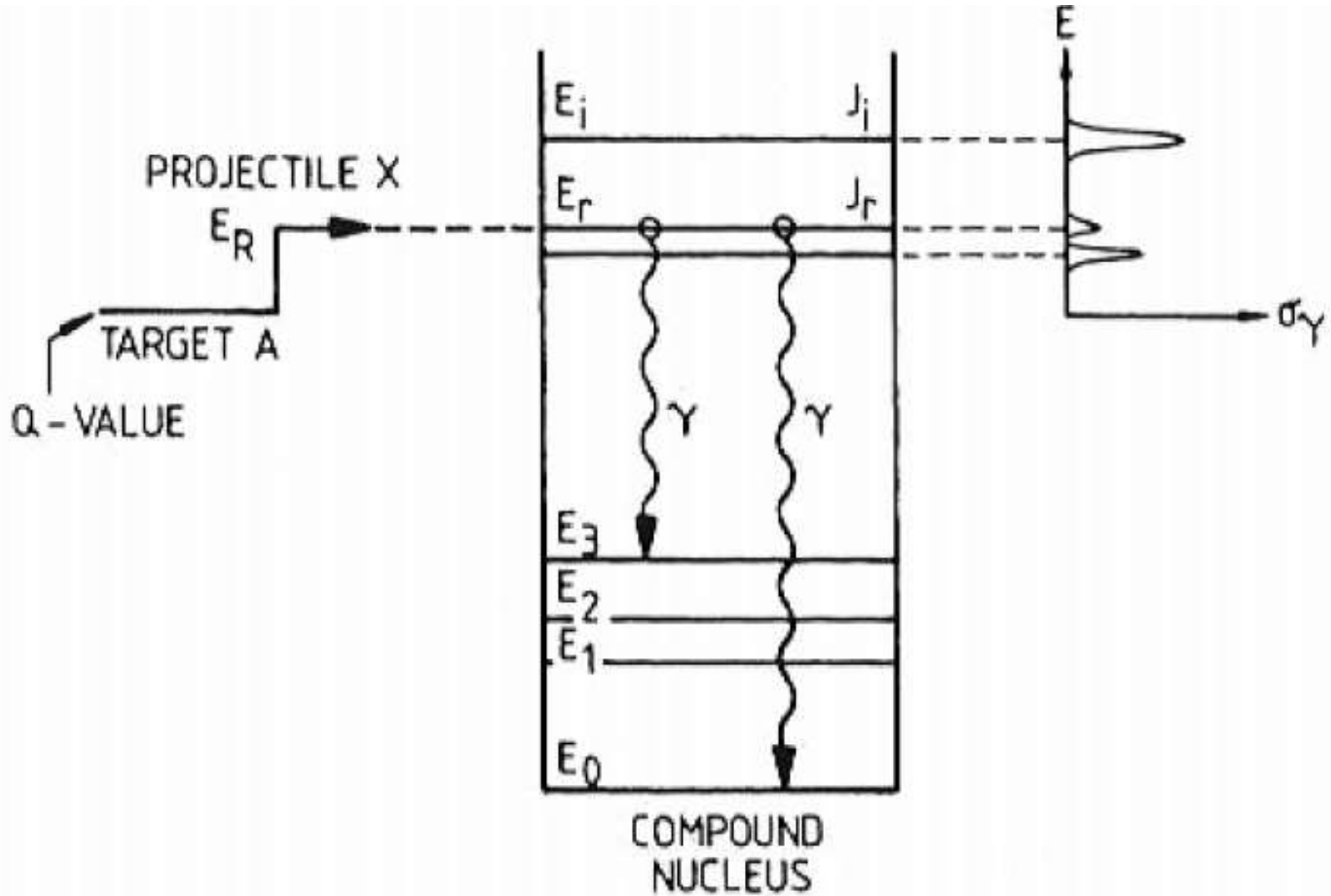


Two separable steps:

1. Incident particle merges with target to form a compound nucleus.  
Intermediate state lives “long” (fs – as)  $\Rightarrow$  thermal equilibrium
2. The compound nucleus deexcites by emitting gammas or particles.

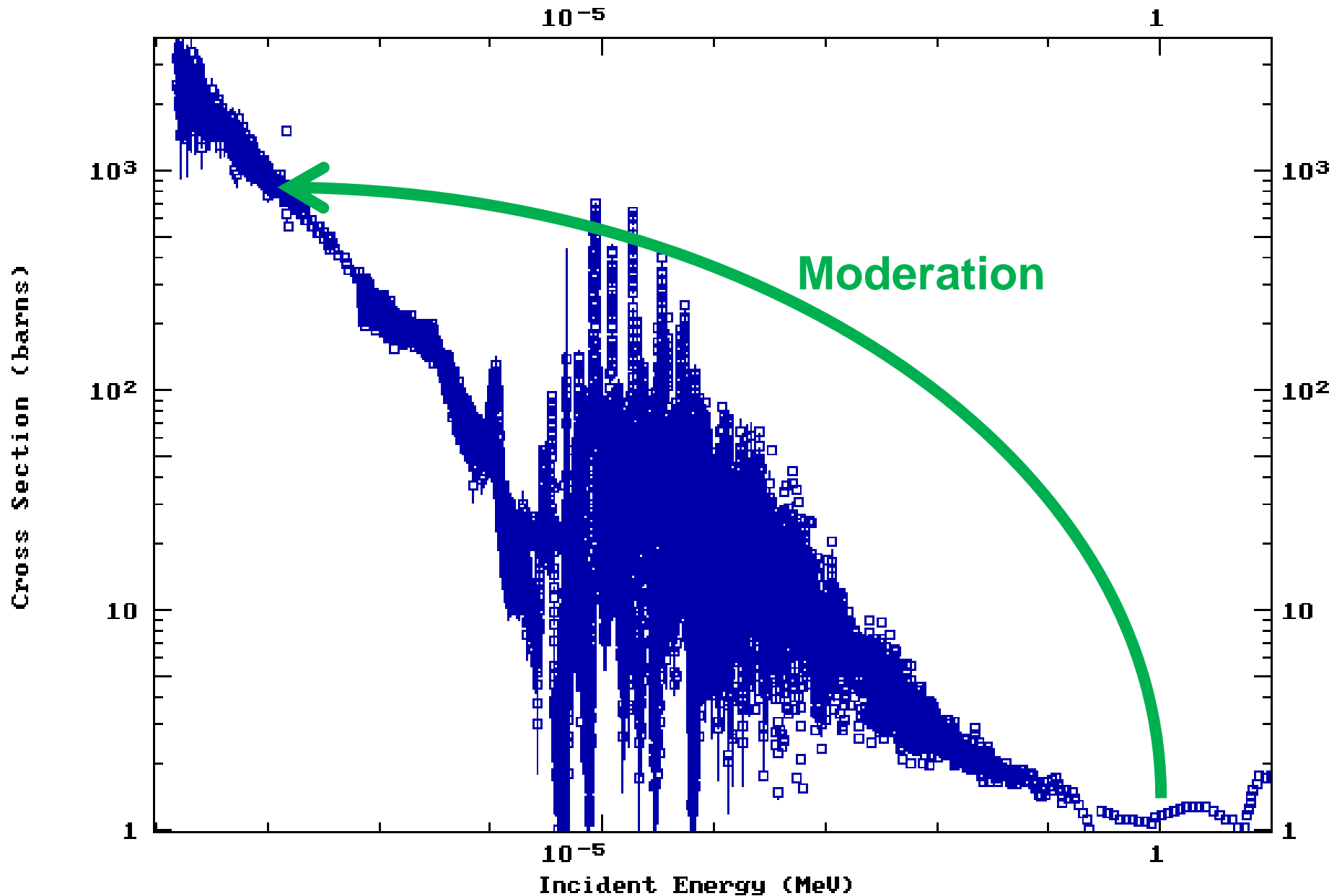
The mode of decay (2.) is independent from the way the compound nucleus was initially formed (1.).

# Resonance reactions

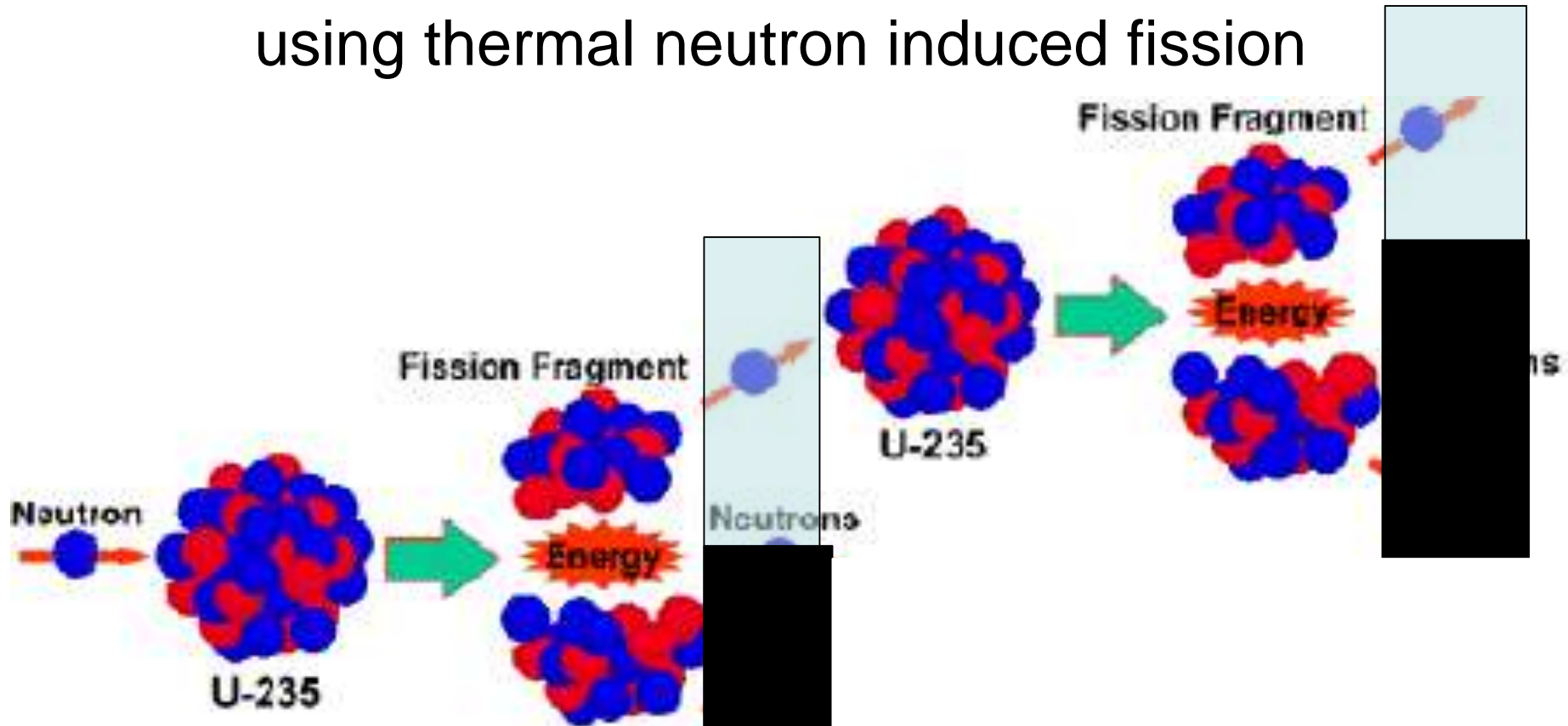


Enhanced cross-section (resonance) when energy matches excited state in compound nucleus.

# $^{235}\text{U}(n,f)$ cross-section as function of energy



# A controlled nuclear chain reaction using thermal neutron induced fission



1. Moderate neutrons
2. Control neutron losses

$k = \text{multiplication factor} = \frac{(\text{neutrons produced in one generation})}{(\text{neutrons produced in previous generation})}$



# Prompt neutron kinetics

Prompt neutron lifetime  $\tau_p$  is the average time between the birth of prompt fission neutrons and their final absorption.

## Assumptions:

- No delayed neutrons
- Infinite reactor, multiplication factor  $k_\infty = k$

time	N(t)
0	n
$\tau_p$	kn
$2\tau_p$	$k^2n$
$3\tau_p$	$k^3n$

$$\frac{dn}{dt} = \frac{k-1}{\tau_p} n \Rightarrow n(t) = n(0) e^{\frac{(k-1)}{\tau_p} t}$$

Time constant

$$T = \frac{\tau_p}{k-1}$$

Exponential decrease ( $k < 1$ )  
or exponential growth ( $k > 1$ )

cf. demographic projections for Germany  
Fertility: 1.5 child/women  $\rightarrow k=0.75$   
 $T=25 \text{ years} / (1-0.75) = 100 \text{ years}$

# Prompt neutron kinetics

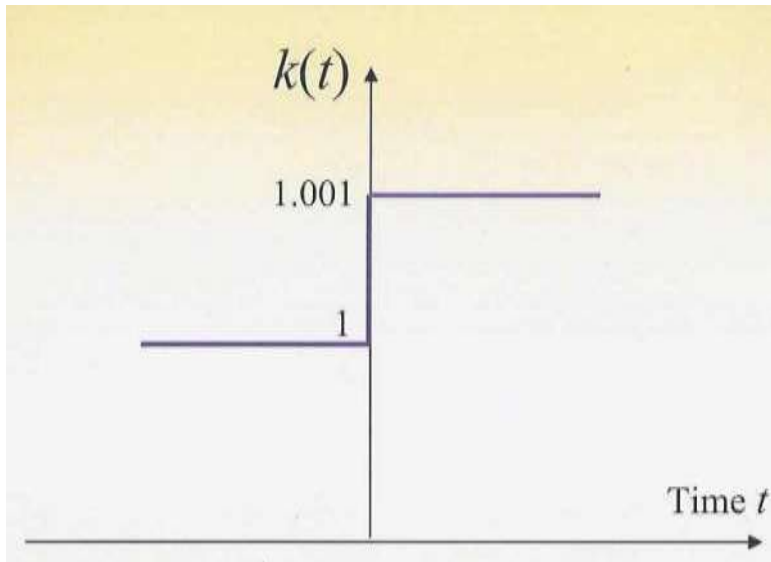
$\tau_p = \tau_s + \tau_d =$  slowing down time + diffusion time

In thermal reactors:  $\tau_s \ll \tau_d$ , i.e.  $\tau_p \cong \tau_d$

$$\tau_d \cong \lambda_a / v \cong 10 \text{ cm} / (2000 \text{ m/s})$$

$$\tau_p \cong \tau_d \cong 50 \text{ } \mu\text{sec}$$

Example: step of reactivity from  $k=1.000$  to  $k=1.001$



$$T = \frac{\tau_p}{k-1} = \frac{50 \cdot 10^{-6}}{10^{-3}} = 0.05 \text{ sec}$$

$$n(t) = n_0 e^{\frac{t}{0.05}}$$

$$\frac{n(1 \text{ sec})}{n_0} = e^{20} = 5E8$$

**“Prompt” control is not possible!**

# Chernobyl: a criticality accident



# An interesting equation

$$n(t) = n(0) \exp((k-1)/\tau_{\text{cycle}} t)$$

$$T_2 = \ln(2) \tau_{\text{cycle}} / (k-1)$$

$k = R_0$  basic reproduction number

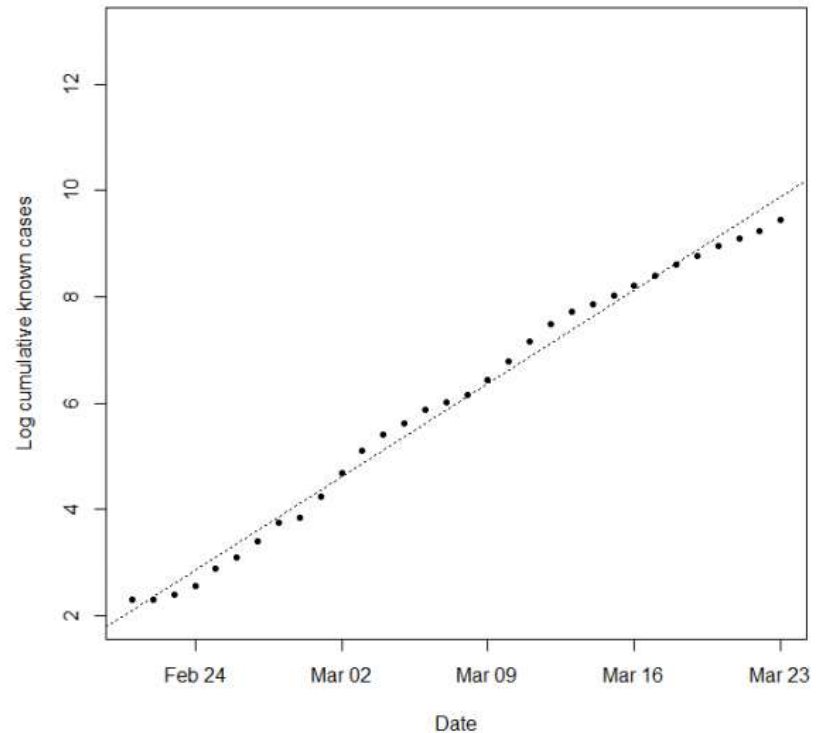
$\tau_{\text{cycle}} \approx$  incubation period

$$k < 1: T_{1/2} = \ln(2) \tau_{\text{cycle}} / (k-1)$$

$$\tau_{\text{cycle}} = 7 \text{ d}; k = 0.8 \Rightarrow T_{1/2} = 3.5 \text{ weeks}$$

$$\tau_{\text{cycle}} = 7 \text{ d}; k = 0.9 \Rightarrow T_{1/2} = 7 \text{ weeks}$$

$$\tau_{\text{cycle}} = 7 \text{ d}; k = 0.99 \Rightarrow T_{1/2} = 70 \text{ weeks}$$



**The greatest shortcoming of the human race is our inability to understand the exponential function. [Prof. AI Bartlett]**



“Just stay calm. It will go away.”



“I continue to shake hands.”

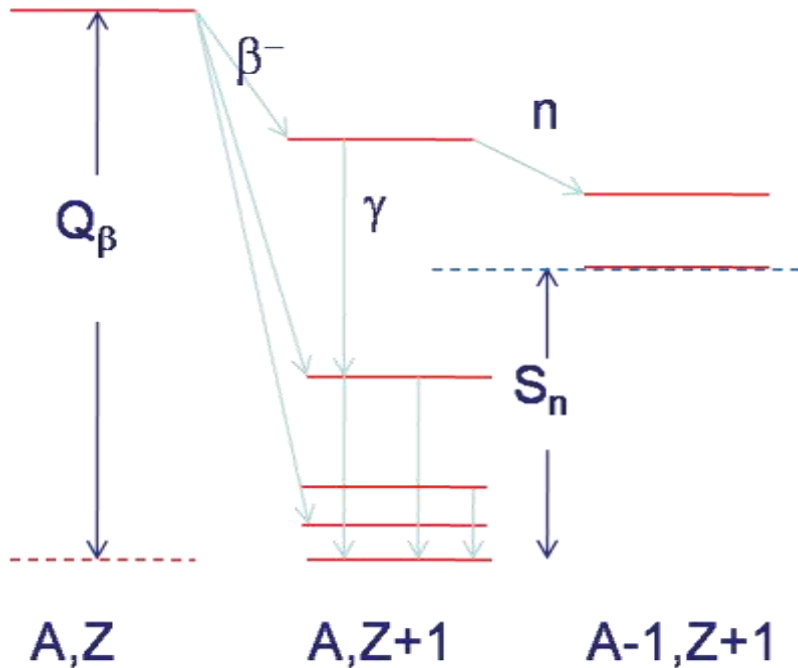


“Everything passes and this will pass.”



“Brazilians never catch anything.”

# Delayed neutron emission from fission products



Possible precursors
$^{87}\text{Br}$
$^{88}\text{Br}$ , $^{137}\text{I}$ , $^{136}\text{Te}$ , $^{134}\text{Sb}$ , $^{141}\text{Cs}$
$^{89}\text{Br}$ , $^{138}\text{I}$ , $^{92,93}\text{Rb}$ , $^{147}\text{La}$ , $^{87}\text{Se}$ , $^{84}\text{As}$
$^{85}\text{As}$ , $^{90}\text{Br}$ , $^{135}\text{Sb}$ , $^{94}\text{Rb}$ , $^{139}\text{I}$ , $^{98,99}\text{Y}$ , $^{142}\text{Cs}$ , $^{80}\text{Ga}$
$^{86,87}\text{As}$ , $^{136}\text{Sb}$ , $^{147,148}\text{Ba}$ , $^{81,82}\text{Ga}$ , $^{140,141}\text{I}$ , $^{91}\text{Br}$ , $^{134}\text{Sn}$ , $^{145}\text{Cs}$ , $^{89}\text{Se}$
$^{83}\text{Ga}$ , $^{146,147}\text{Cs}$ , $^{95,96,97,98,99}\text{Rb}$ , $^{92}\text{Br}$ , $^{91}\text{Se}$

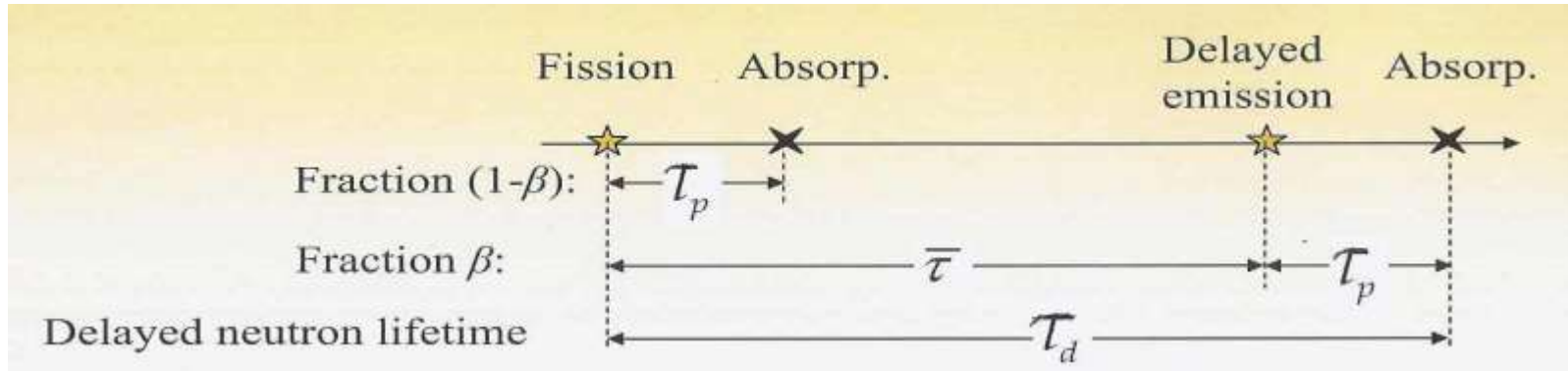
$\beta$  = percentage of delayed neutrons

average emission time

$^{235}\text{U}$	$\beta$	$\nu$	$\nu_p$	$\nu_d$
Ther	0.0067	2.490	2.473	0.01668

$$\tau_{\text{delayed}} = \frac{1}{\beta} \sum_i \beta_i \tau_i \approx 12 \text{ sec}$$

# Neutron lifetime, taking into account delayed neutrons



$$k = k_{prompt} + k_{delayed} = 1 = (1 - \beta) + \beta$$

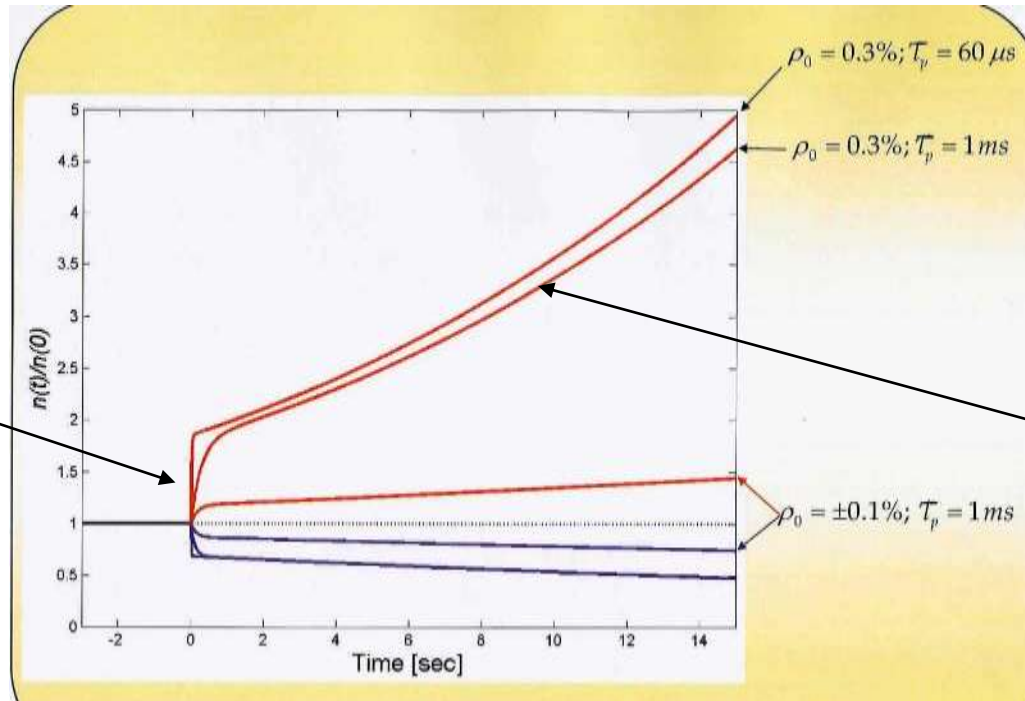
$$\tau = (1 - \beta)\tau_p + \beta(\tau_{delayed} + \tau_p) \approx \beta\tau_{delayed} = 0.08 \text{ sec}$$

Now for step from  $k=1.000$  to  $k=1.001$

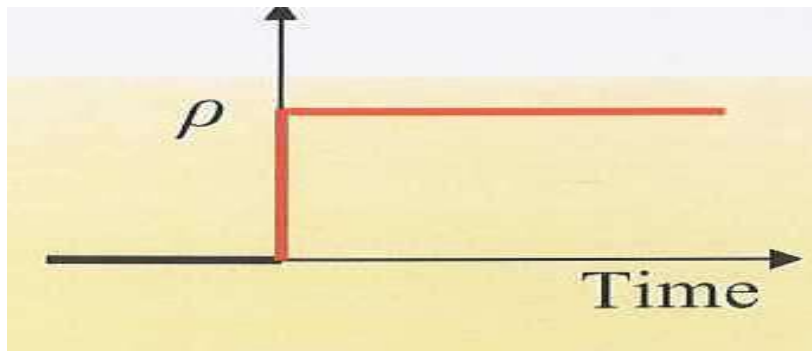
$$T = \beta\tau_{delayed} / (k-1) = 80 \text{ seconds}$$

# Reactor response to a step of reactivity

Prompt jump

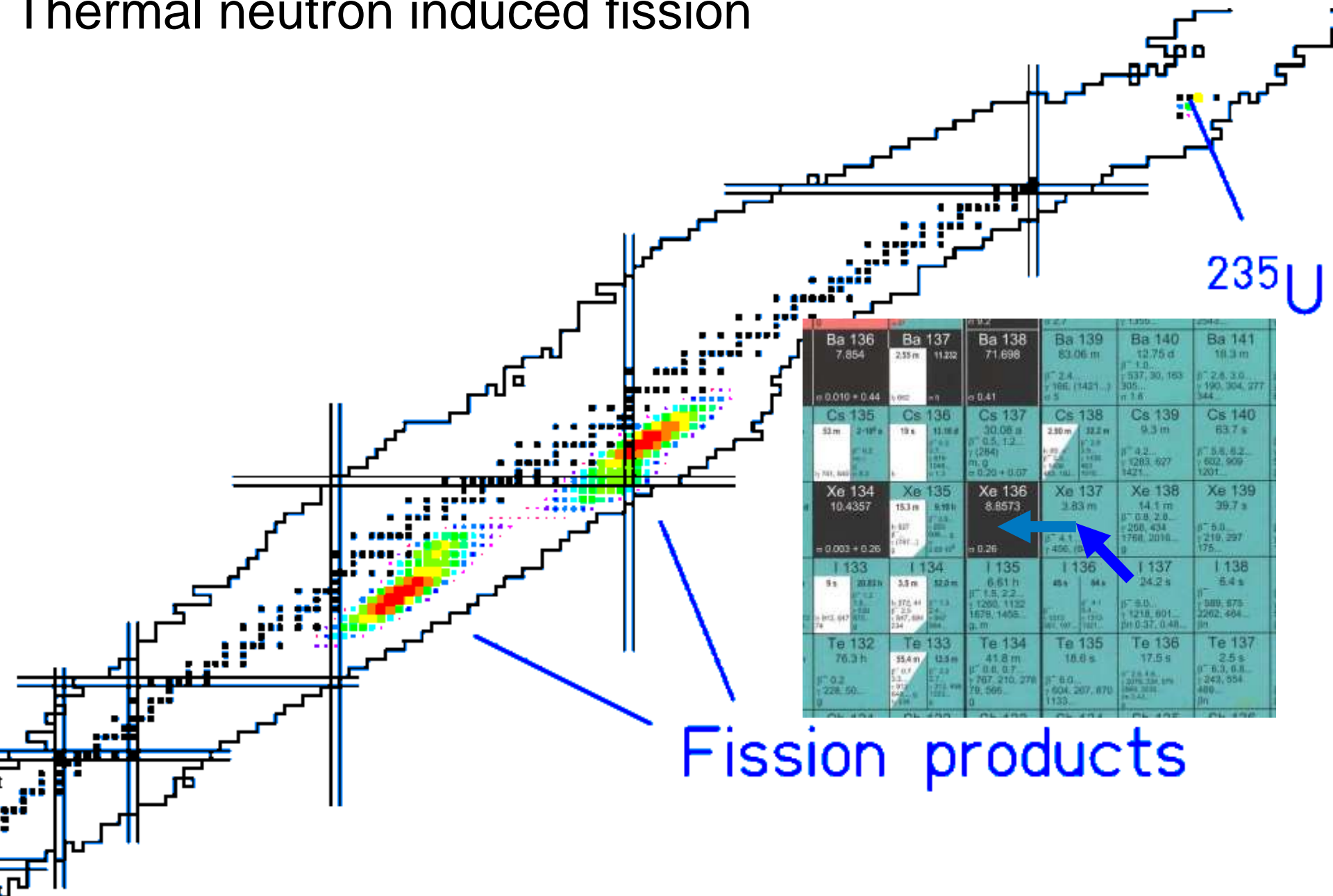


Slow increase due to Delayed neutrons

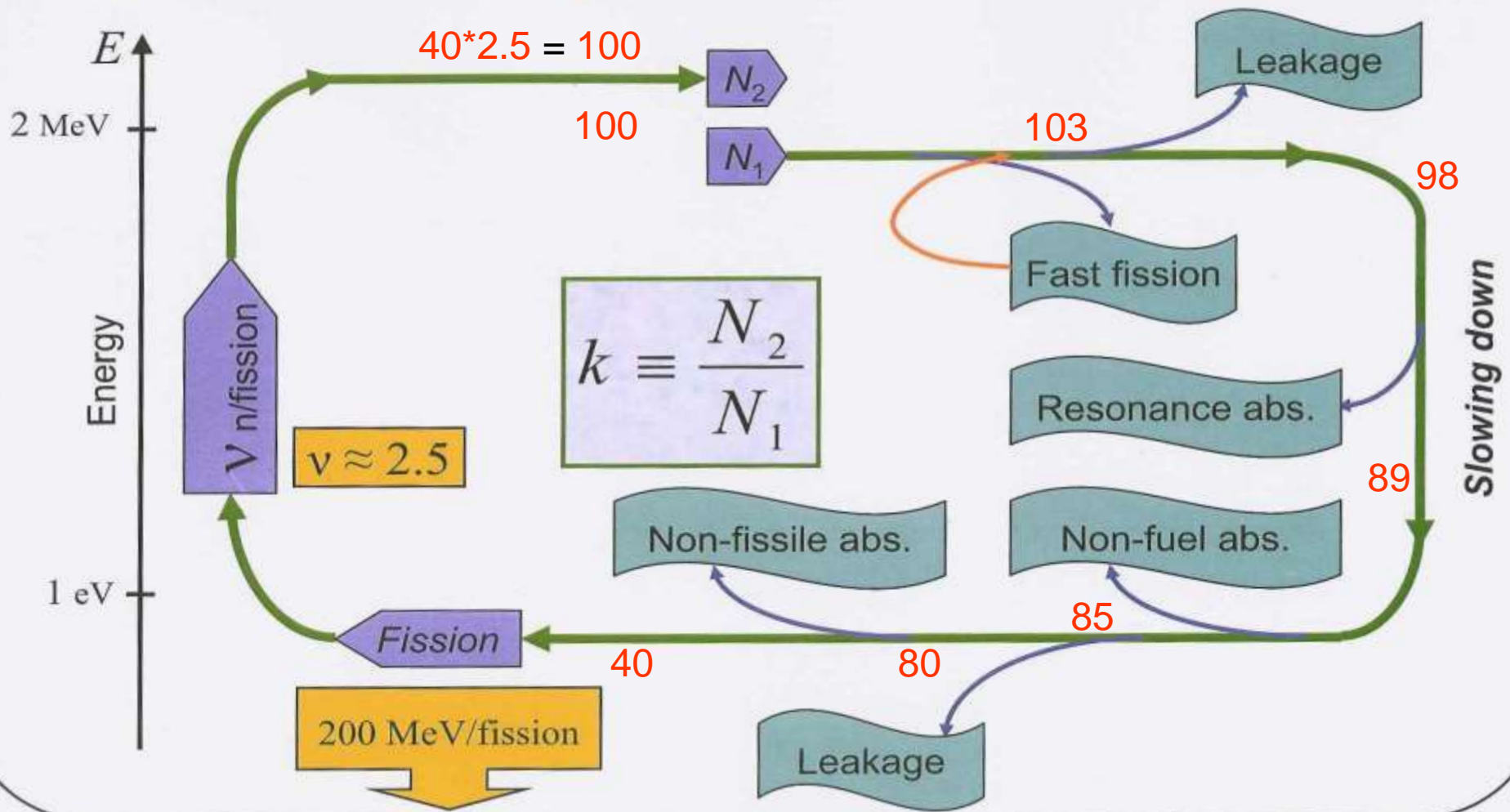




# Thermal neutron induced fission

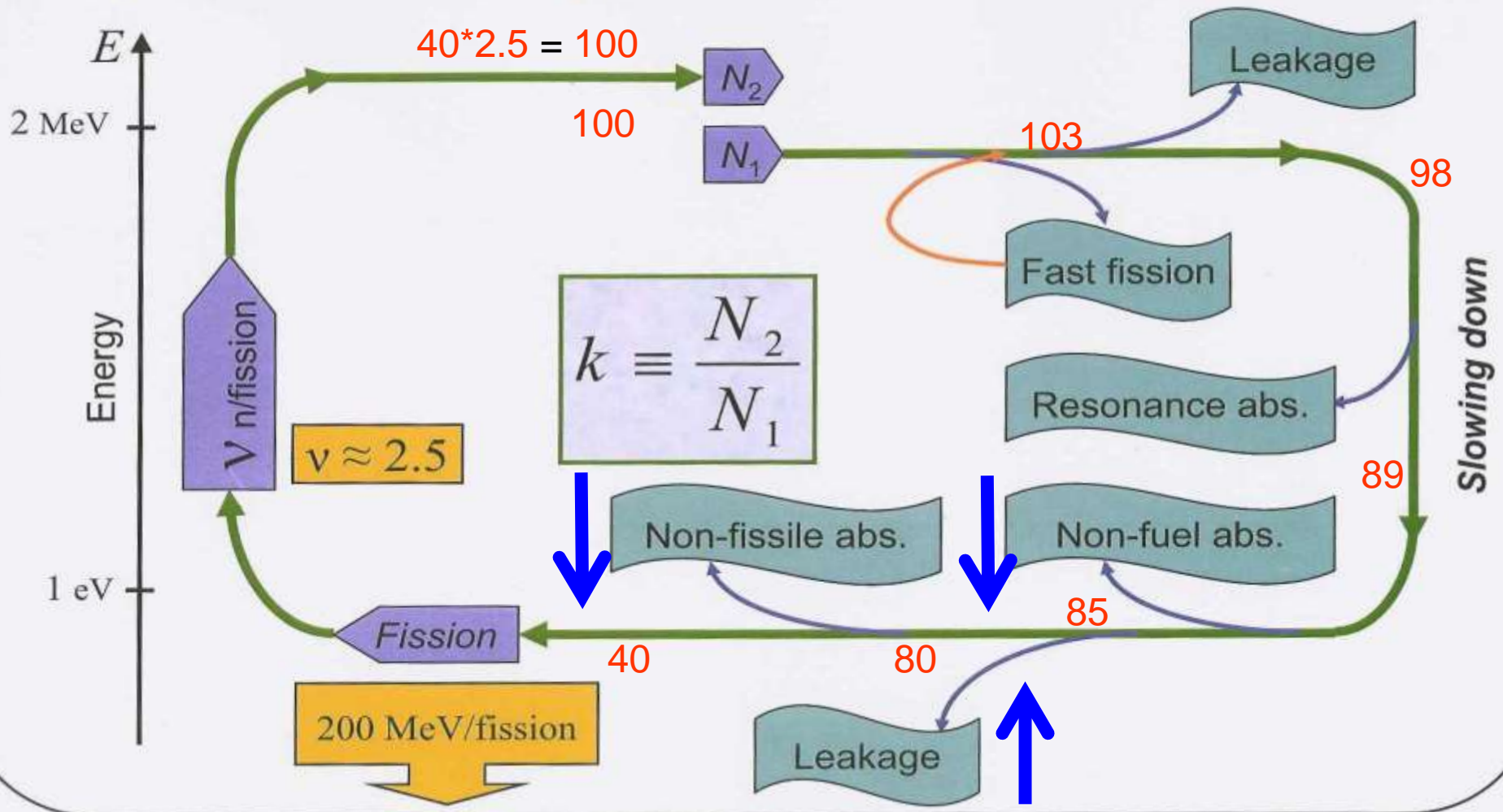


# Physical Principles of a Nuclear Reactor



neutron numbers are given for a typical PWR reactor

# Physical Principles of a Nuclear Reactor



**Research reactor**

# Components of a nuclear reactor

1. Fuel
2. Moderator
3. Control rods
4. Coolant
5. Pressure vessel
6. Containment
7. Steam generator (for power plants) or  
experimental facilities (for research reactors)

# Moderator

elastic collisions with light atoms (mass A):

average energy loss  $E_{n+1} - E_n = 2 E_n A/(A+1)^2$

$$\ln(E_n) - \ln(E_{n+1}) = \xi = 1 - (A-1)^2/(2A) * \ln[(A+1)/(A-1)]$$

$$\Sigma = n \sigma = m/M \sigma$$

Moderating power:

$\xi \Sigma_{\text{scatter}}$

Moderating ratio:

$\xi \Sigma_{\text{scatter}} / \Sigma_{\text{abs.}}$

Light water (H<sub>2</sub>O)

1.28

58

Heavy water (D<sub>2</sub>O)

0.18

21000

Beryllium (Be)

0.16

130

Graphite (C)

0.064

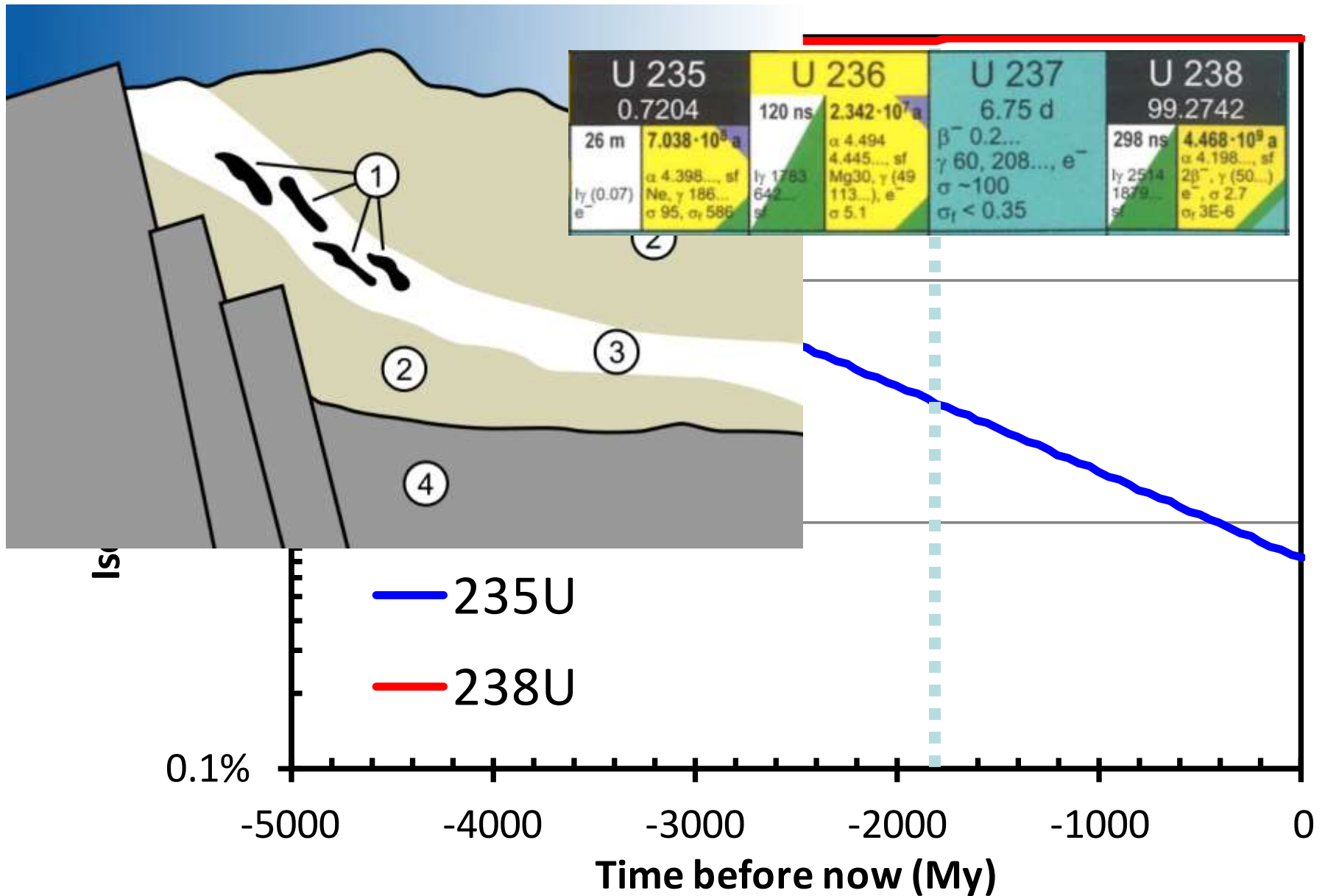
200

Polyethylene (CH<sub>2</sub>)<sub>x</sub>

3.26

122

# Q4: The moderator of the first nuclear reactor



# Choice of coolant

coolant = moderator

⇒ passive regulation

⇒ intrinsic safety

RBMK:

graphite moderator

water cooling

⇒ **positive void  
coefficient !**



# RHF fuel element



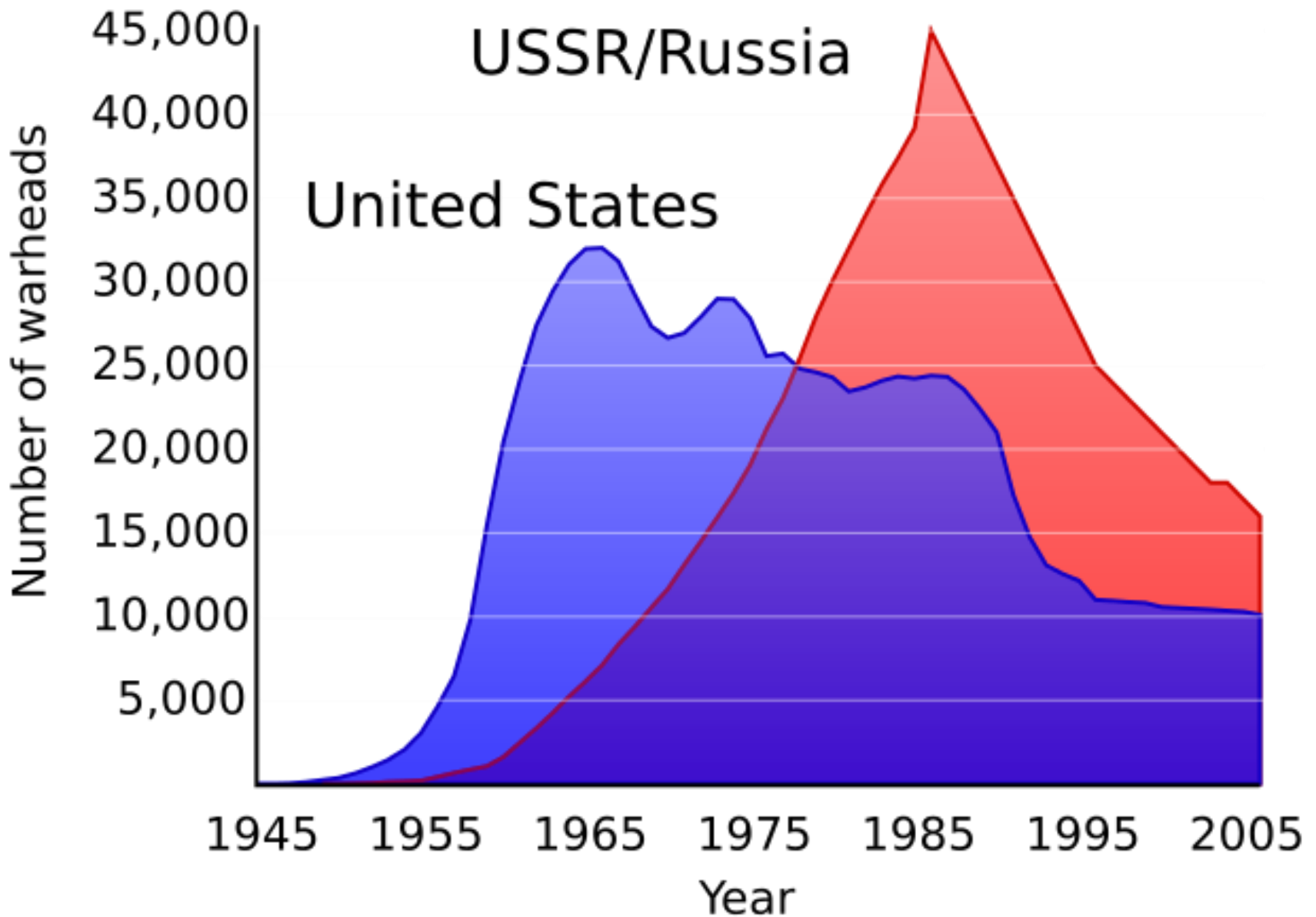
8.6 kg  $^{235}\text{U}$ , 93% enriched





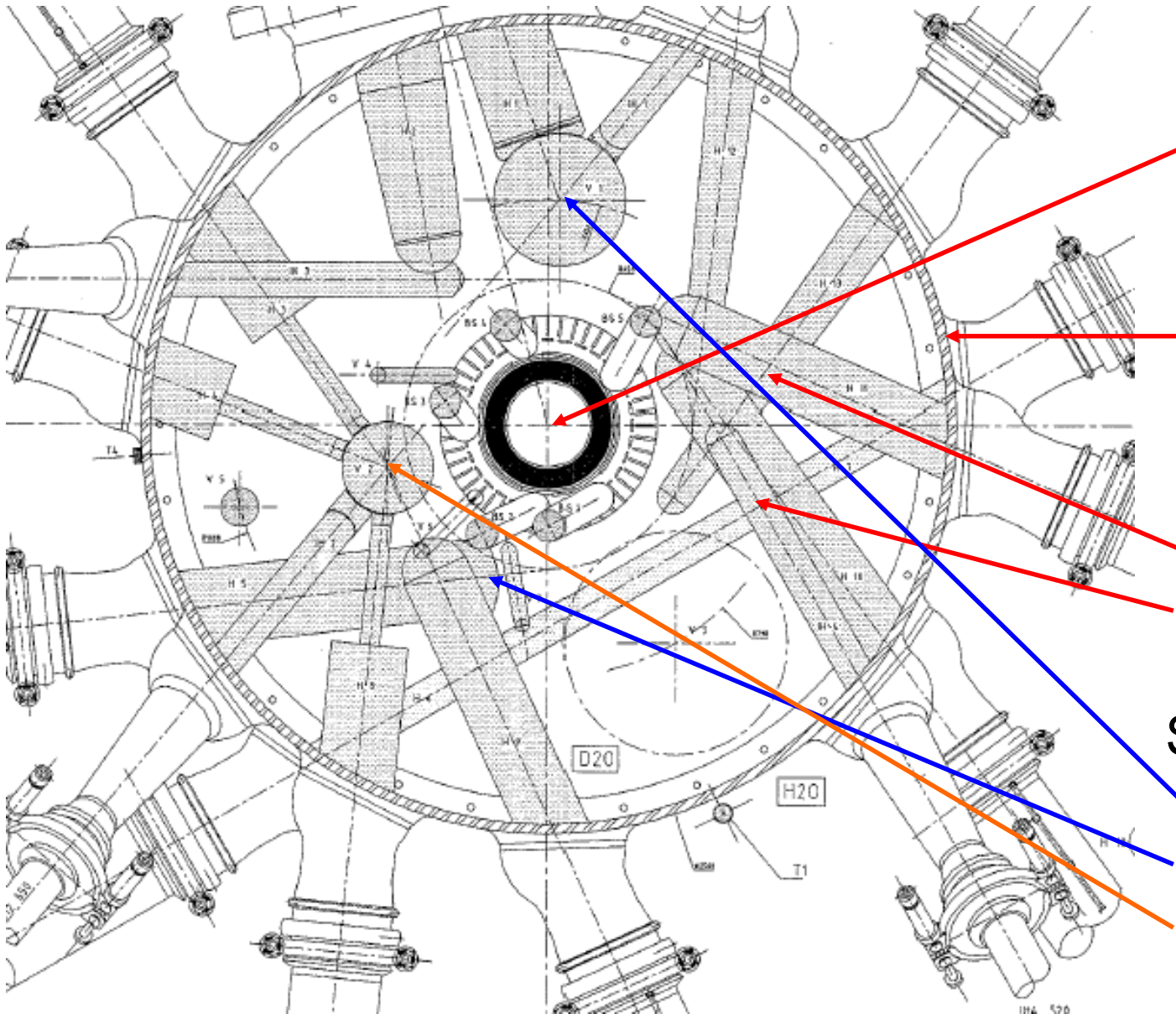
# 8 December 1987: Intermediate-Range Nuclear Forces Treaty





1 warhead = 25 kg HEU = 3 fuel elements for ILL  
The ILL reactor contributes to permanent disarmament!

# The reactor core and vessel



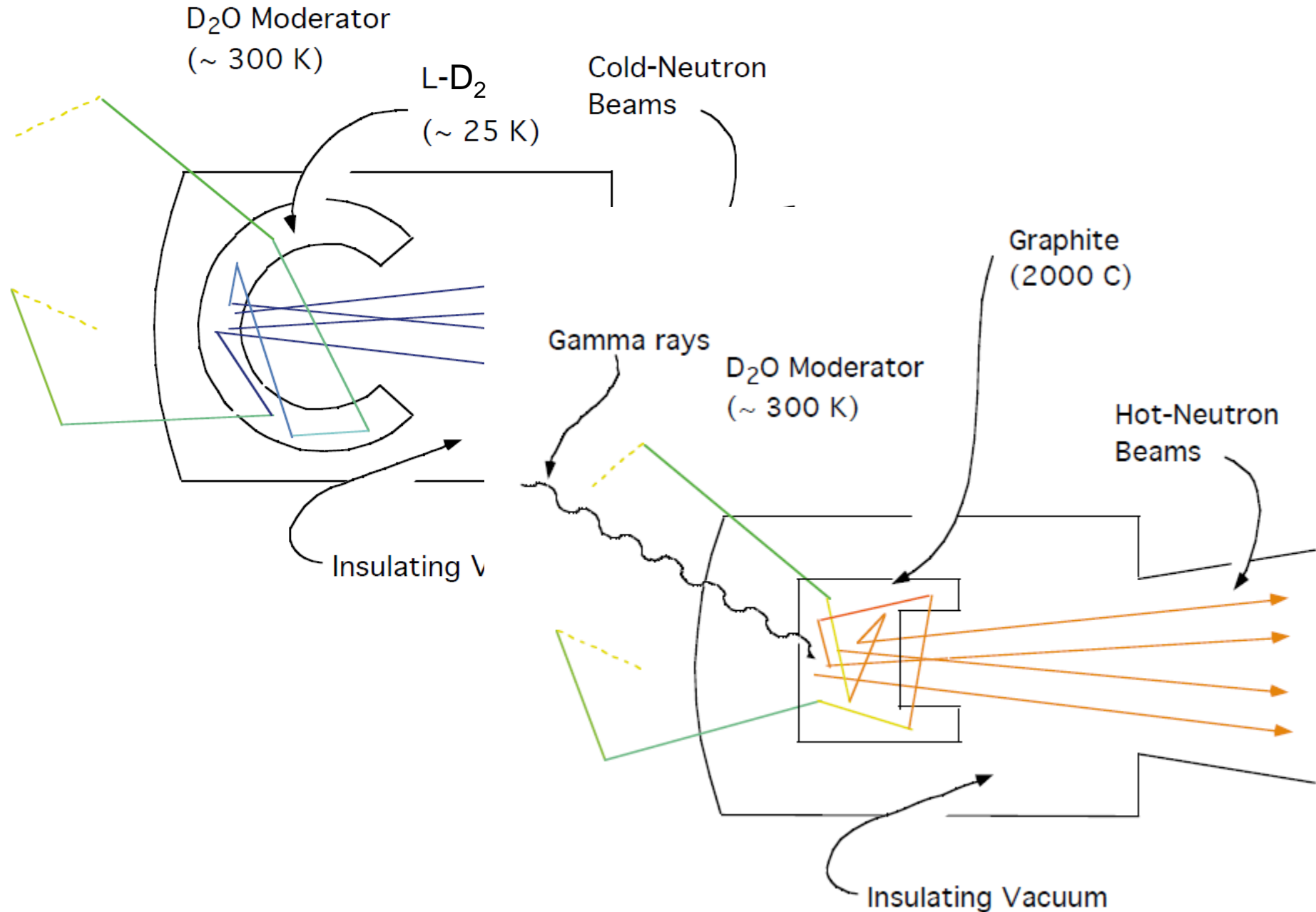
Fuel element :  
-  $R_i = 14$  cm  
-  $R_e = 19$  cm

Vessel:  
-  $R = 125$  cm

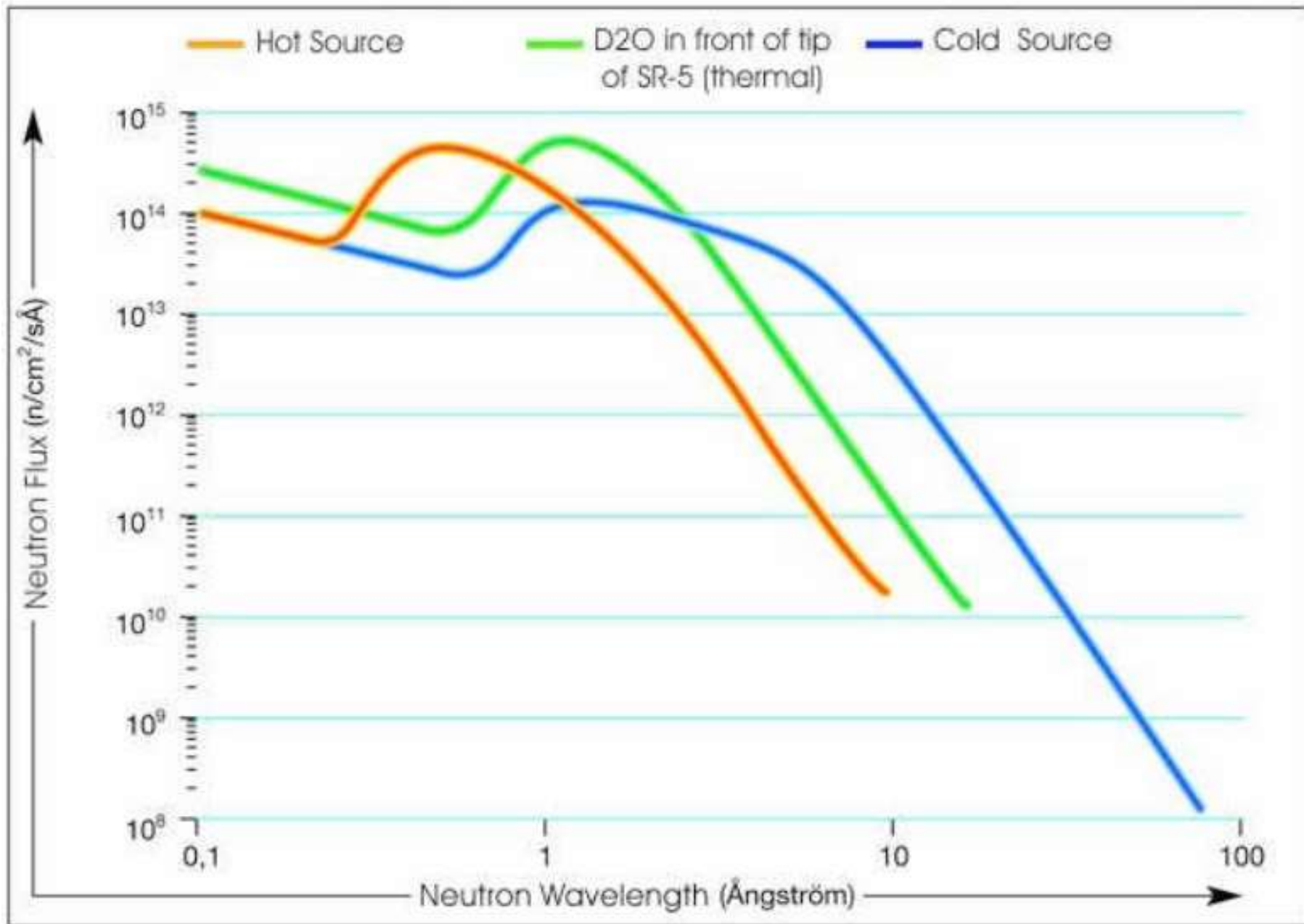
Beam tubes :  
- 13 Horizontal  
- 4 inclined

Sources  
- VCS  
- HCS  
- HS

# Spectral shaping with dedicated moderators



# Spectral shaping with dedicated moderators



## Power reactor

- heat used to produce electricity
- neutrons just to maintain chain reaction
- needs **high power**, **high temperature** and **high pressure** for good thermal efficiency
- BWR: 75 bar, 285°C
- PWR: 155 bar, 315°C
- 25 cm thick steel pressure vessel  $\Rightarrow$  defines lifetime (40..60 y)



## Research reactor

- neutrons used for applications
- heat not used
- operates at **lower power**, **low temperature** (ILL 30-48°C) and **low pressure** (<14 bar)
- vessel and all inserts made from pure Al-alloy
- modular and exchangeable  $\Rightarrow$  no finite lifetime

# ILL: Replacement of the reactor vessel 1990-94

